



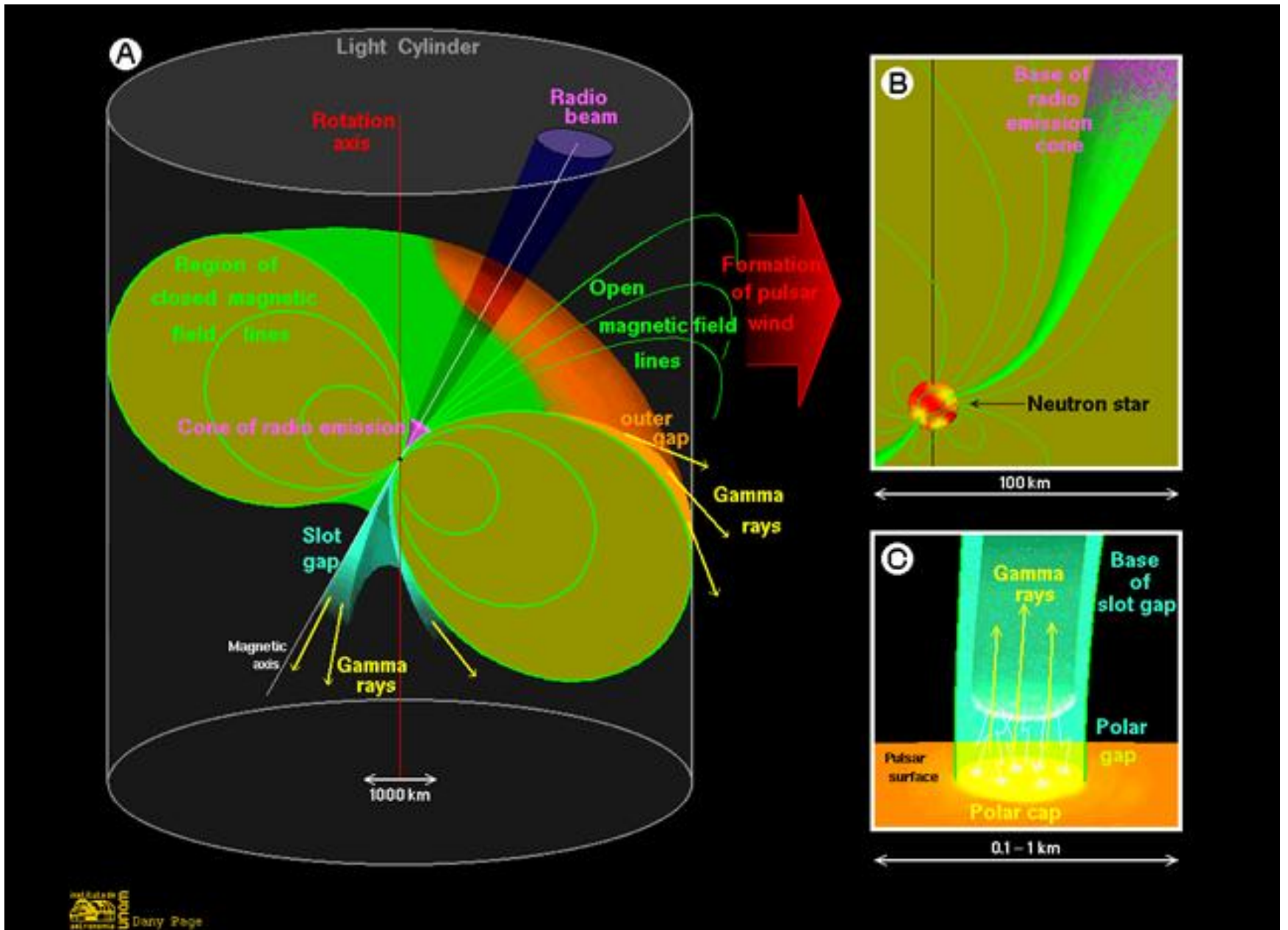
Mapping the B -Field Structure of Elongated PWNe

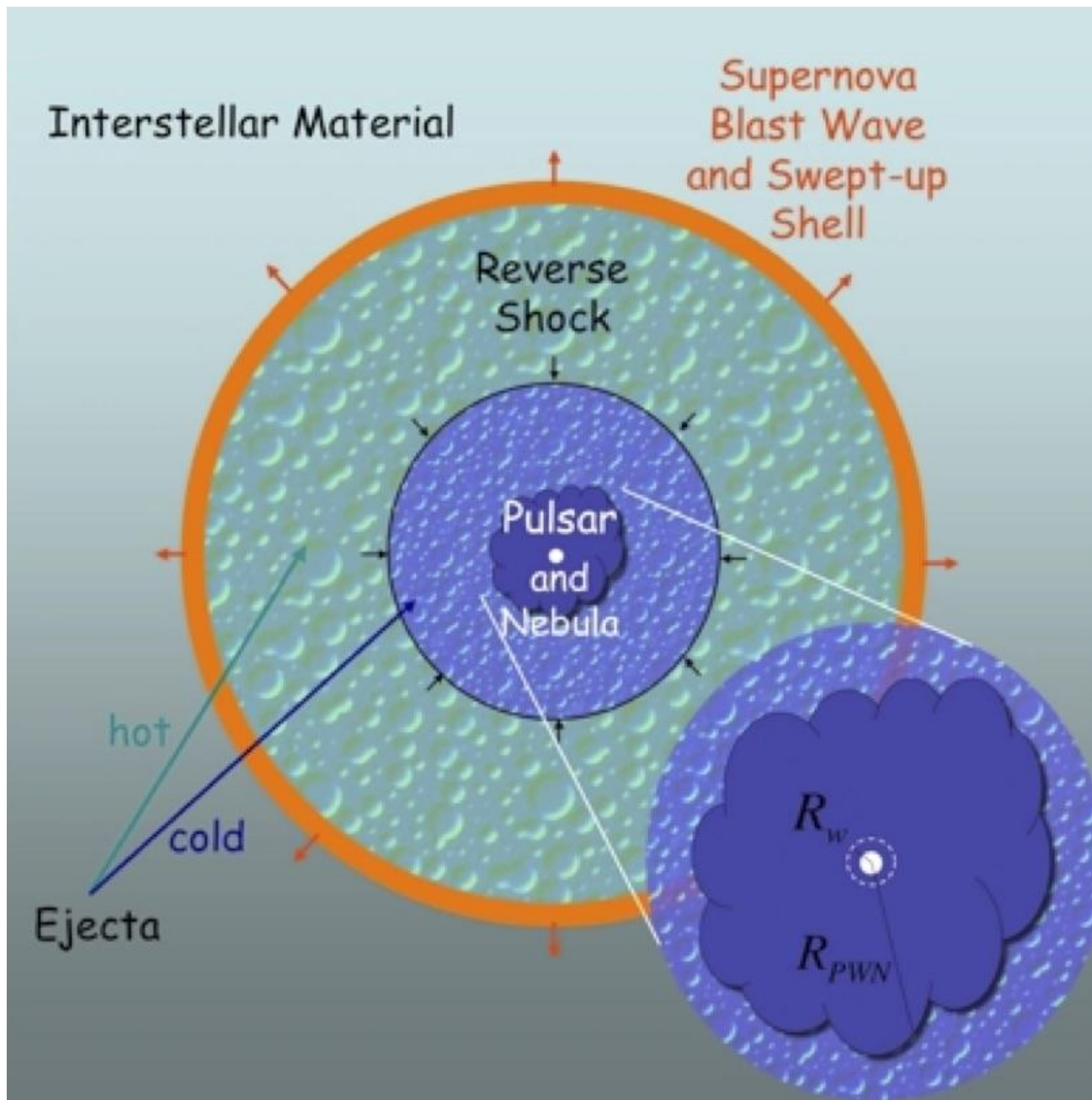
Stephen C.-Y. Ng

HKU

Collaborators

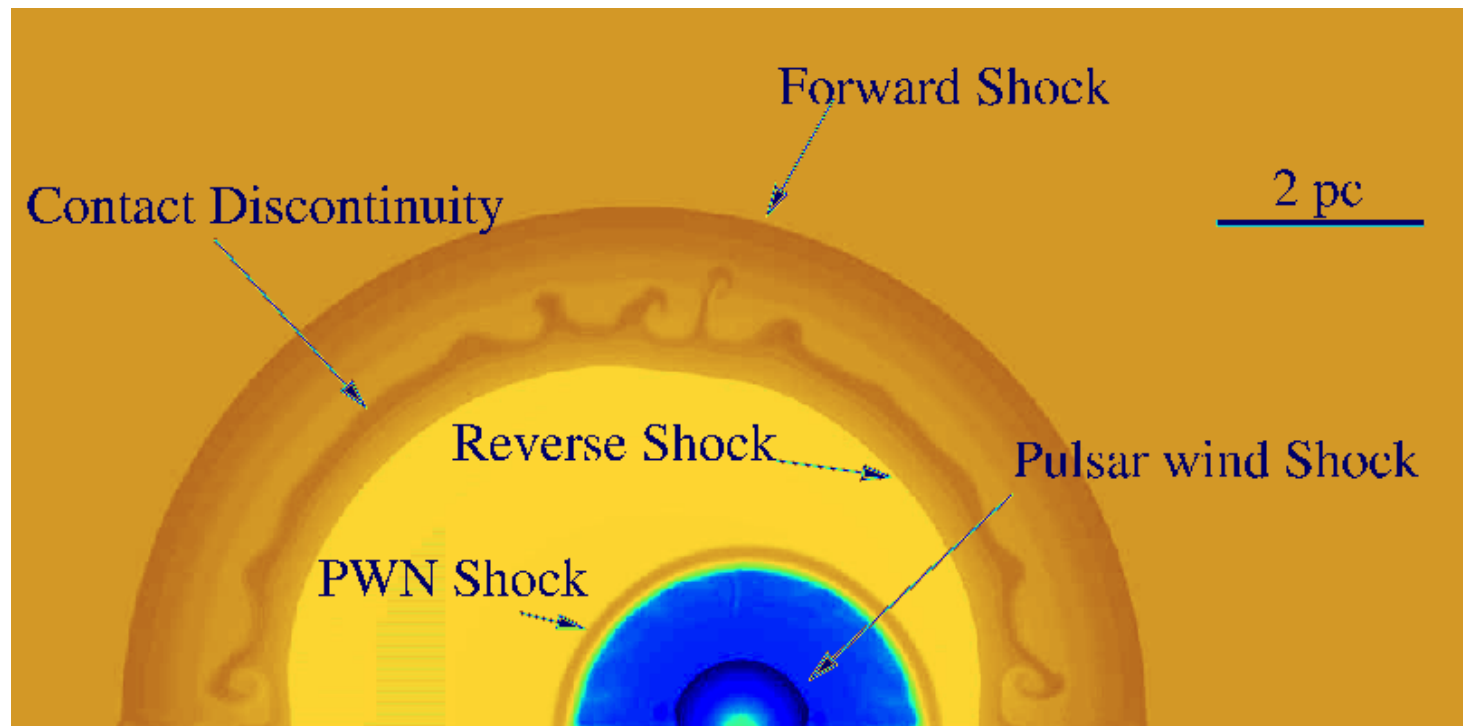
- Y. K. Ma (student)
- W. Y. Leung (student)
- R. Wen (student)
- P. O. Slane
- N. Bucciantini
- T. Temim
- B. M. Gaensler
- Z. Wang





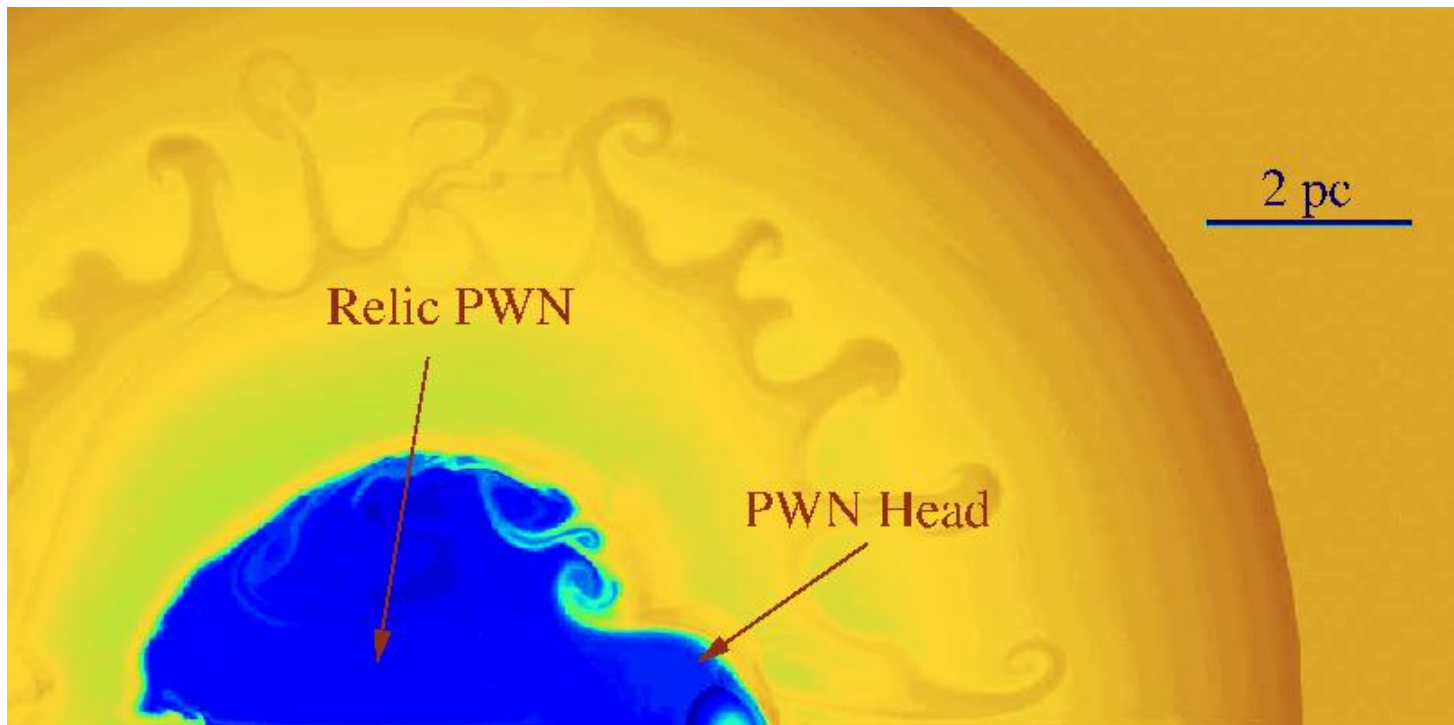
PWN Evolution

- Free expansion phase
- Reverse shock interaction
- Bow shocks



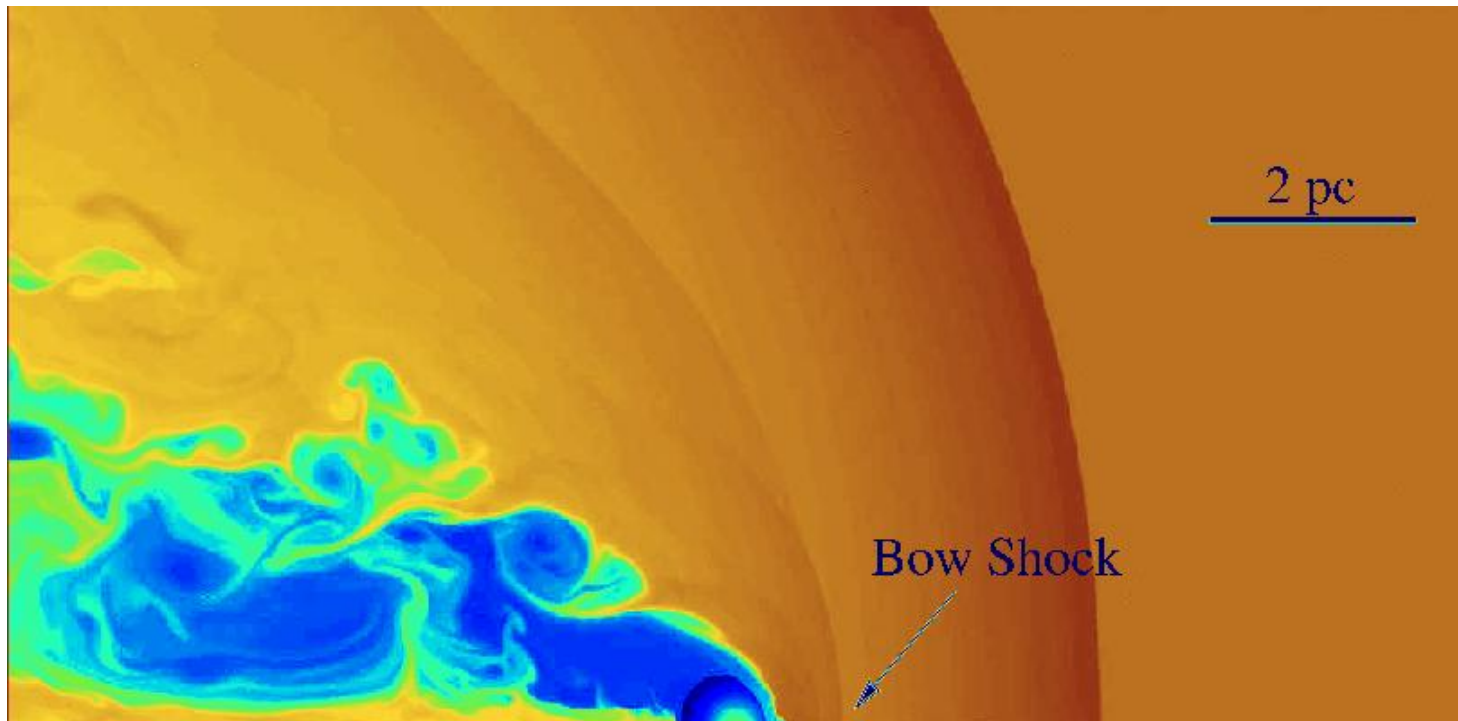
PWN Evolution

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PWN Evolution

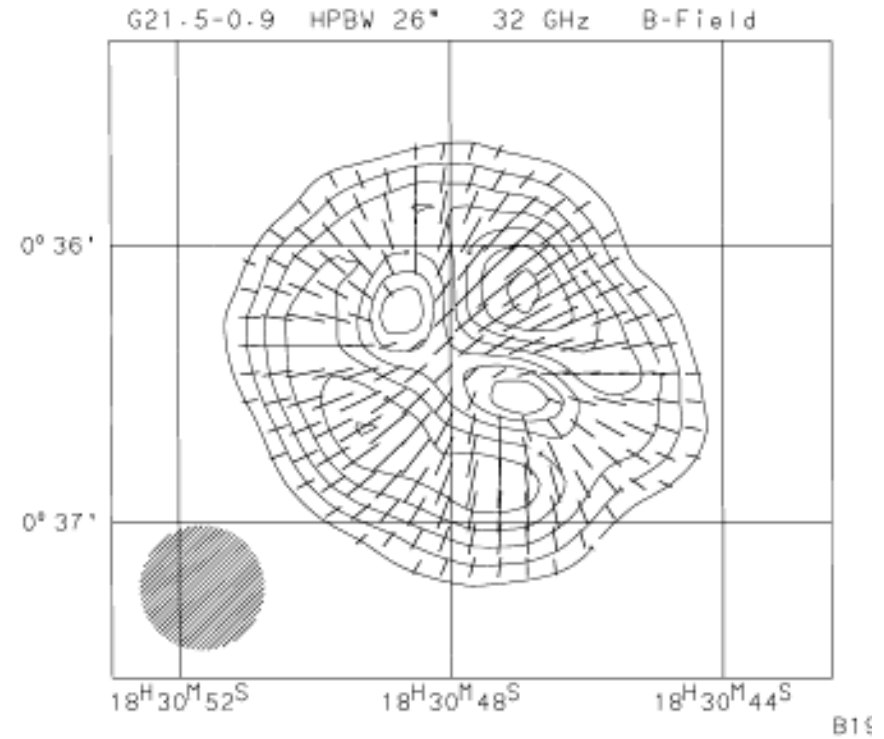
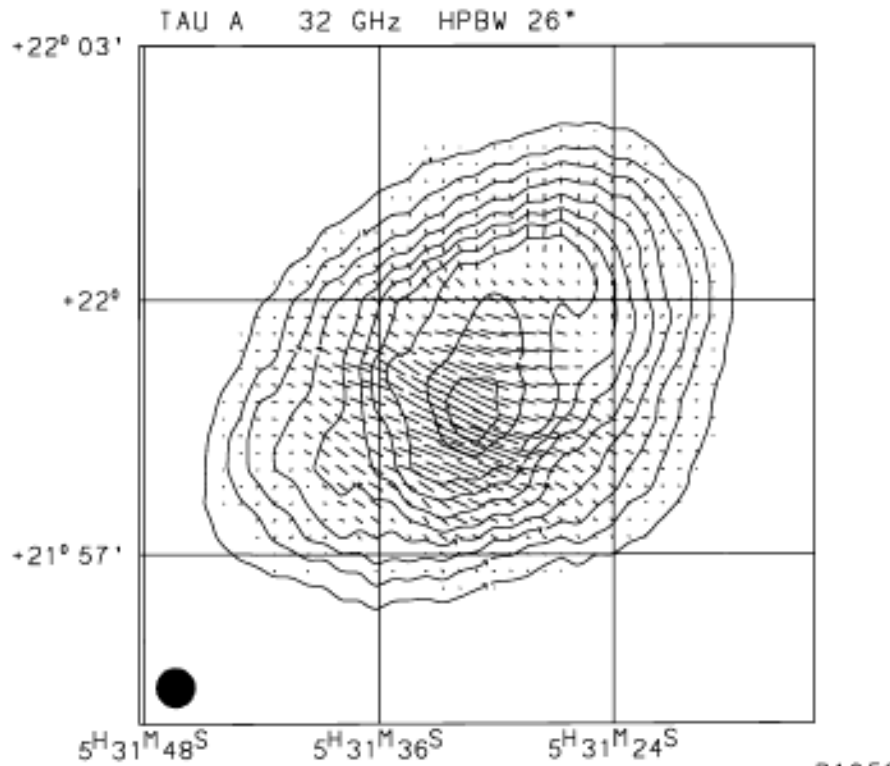
- Free expansion phase
- Reverse shock interaction
- Bow shocks



Important Questions

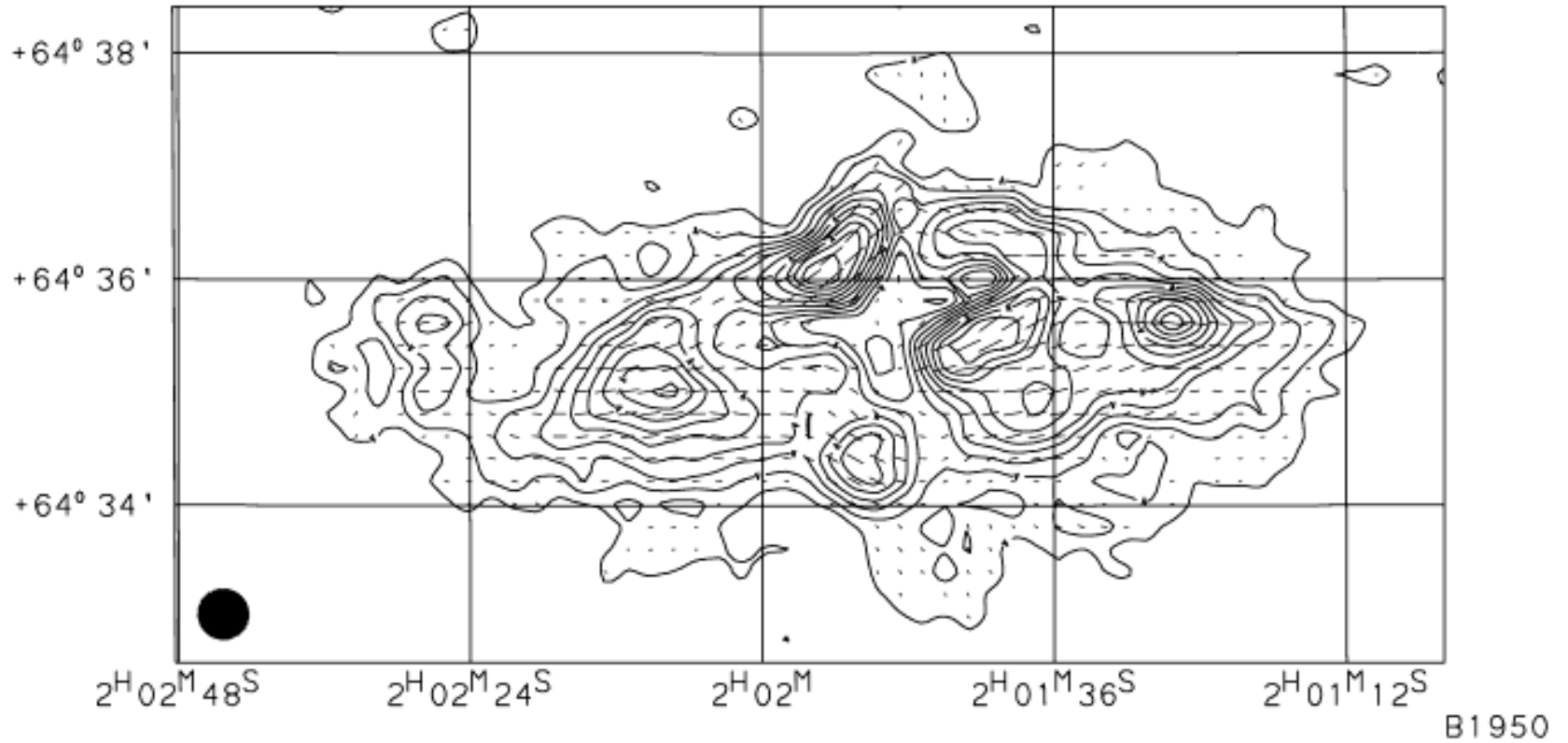
- What is the B -field structure of PWNe?
- What determines the B -field configurations?
- Connection with particle acceleration and transport?

Previous Studies

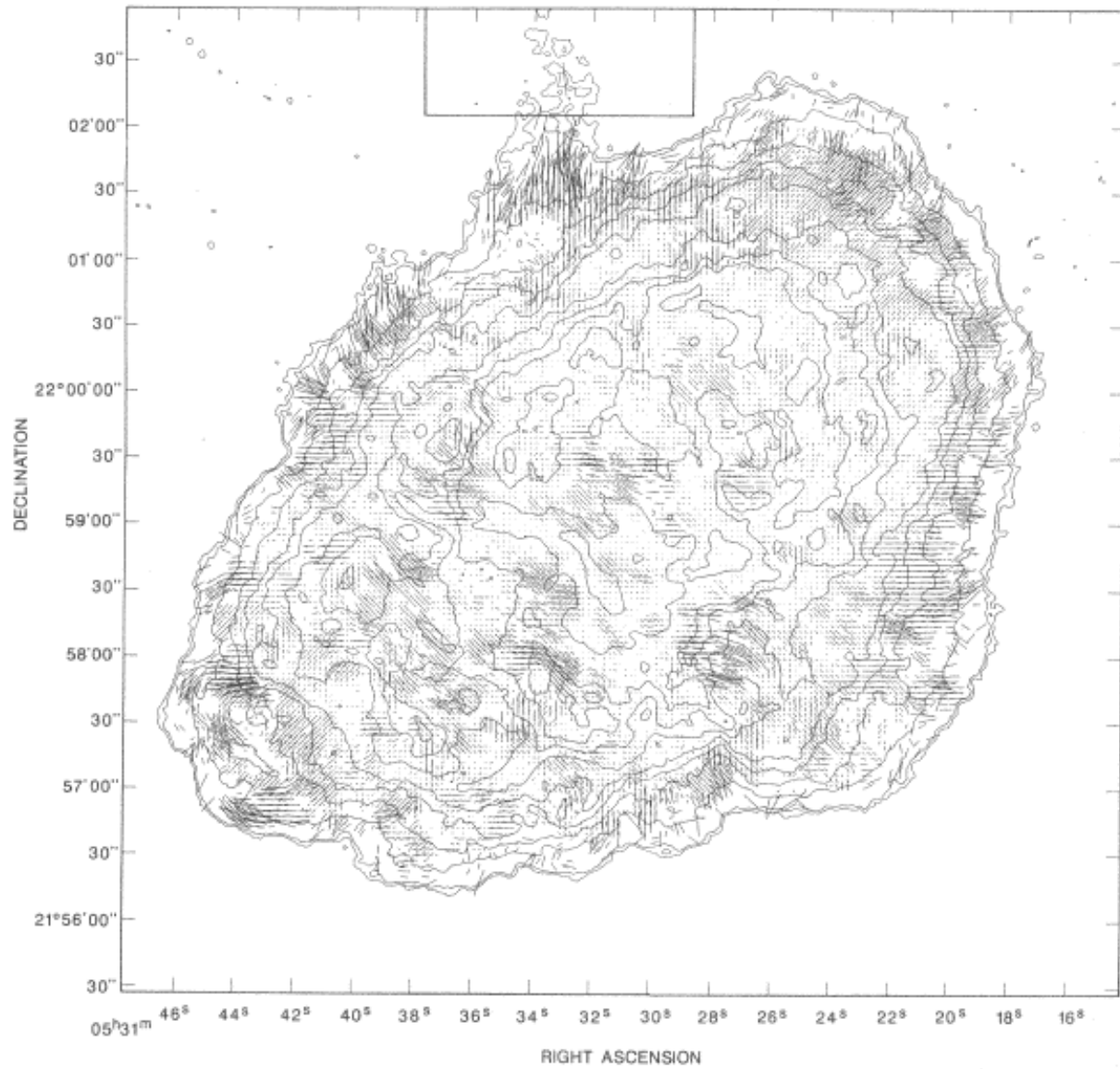


Reich (2002)

3C58 PI HPBW 26" 32 GHz B-Field



Reich (2002)



Bietenholz & Kronberg (1990)

ATCA



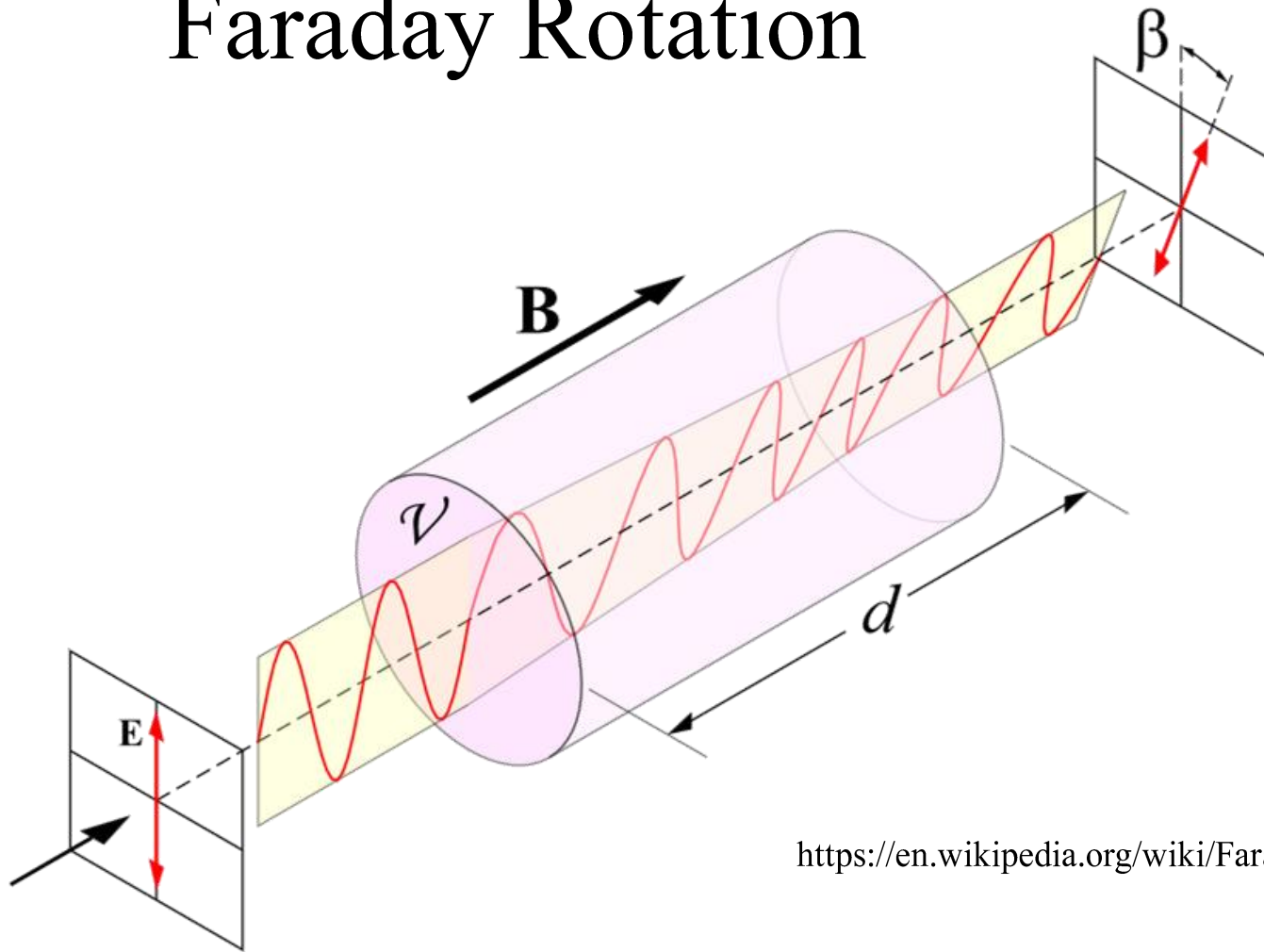
ATNF/CSIRO



HEPRO V

W. Grammer, NRAO/AUI/NSF

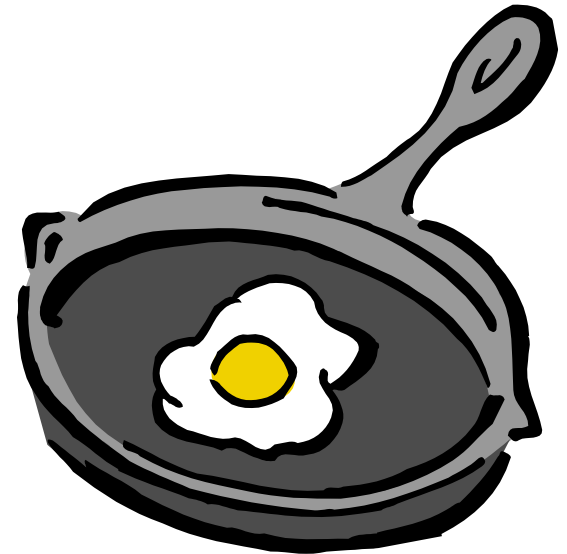
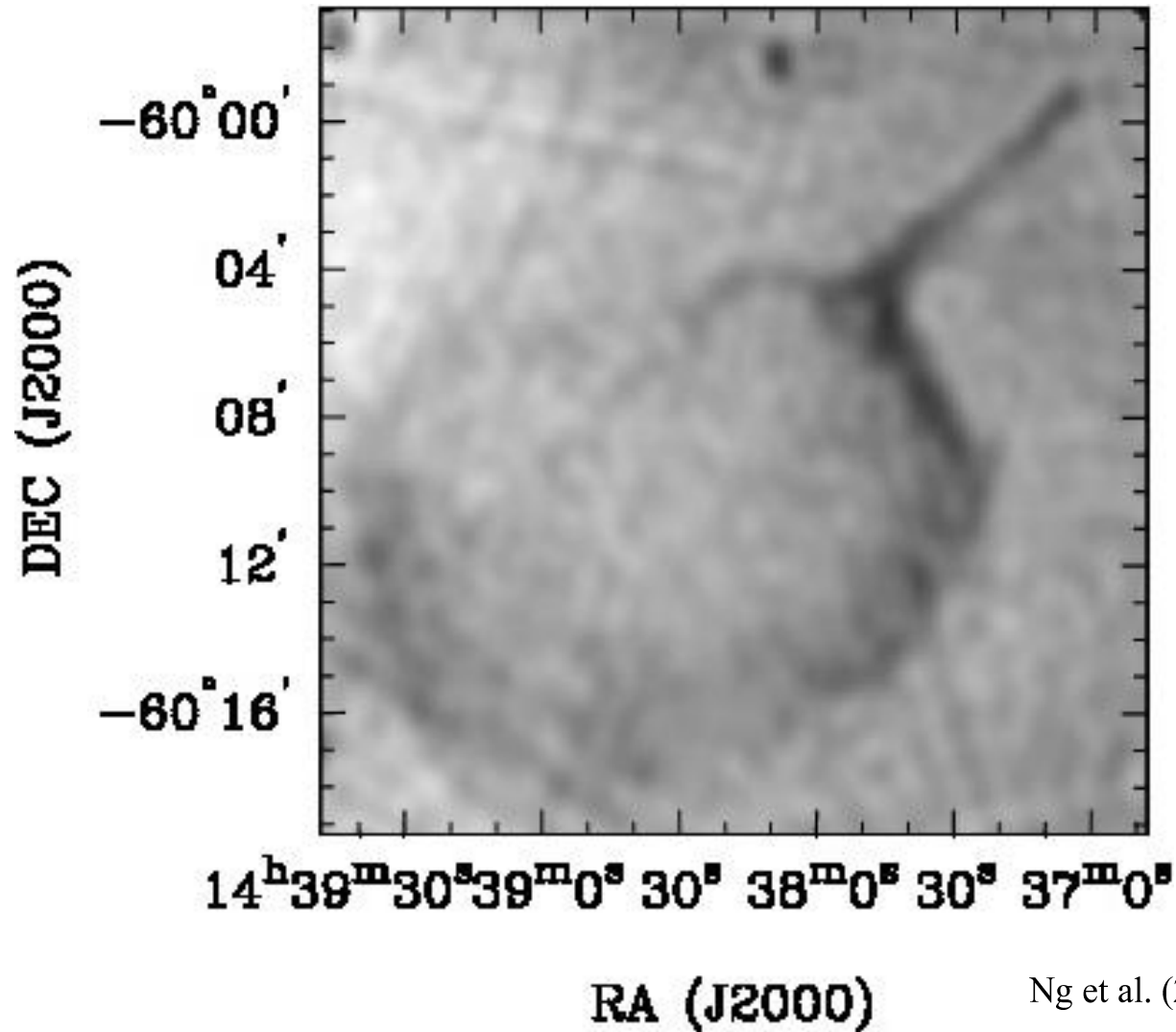
Faraday Rotation



https://en.wikipedia.org/wiki/Faraday_effect

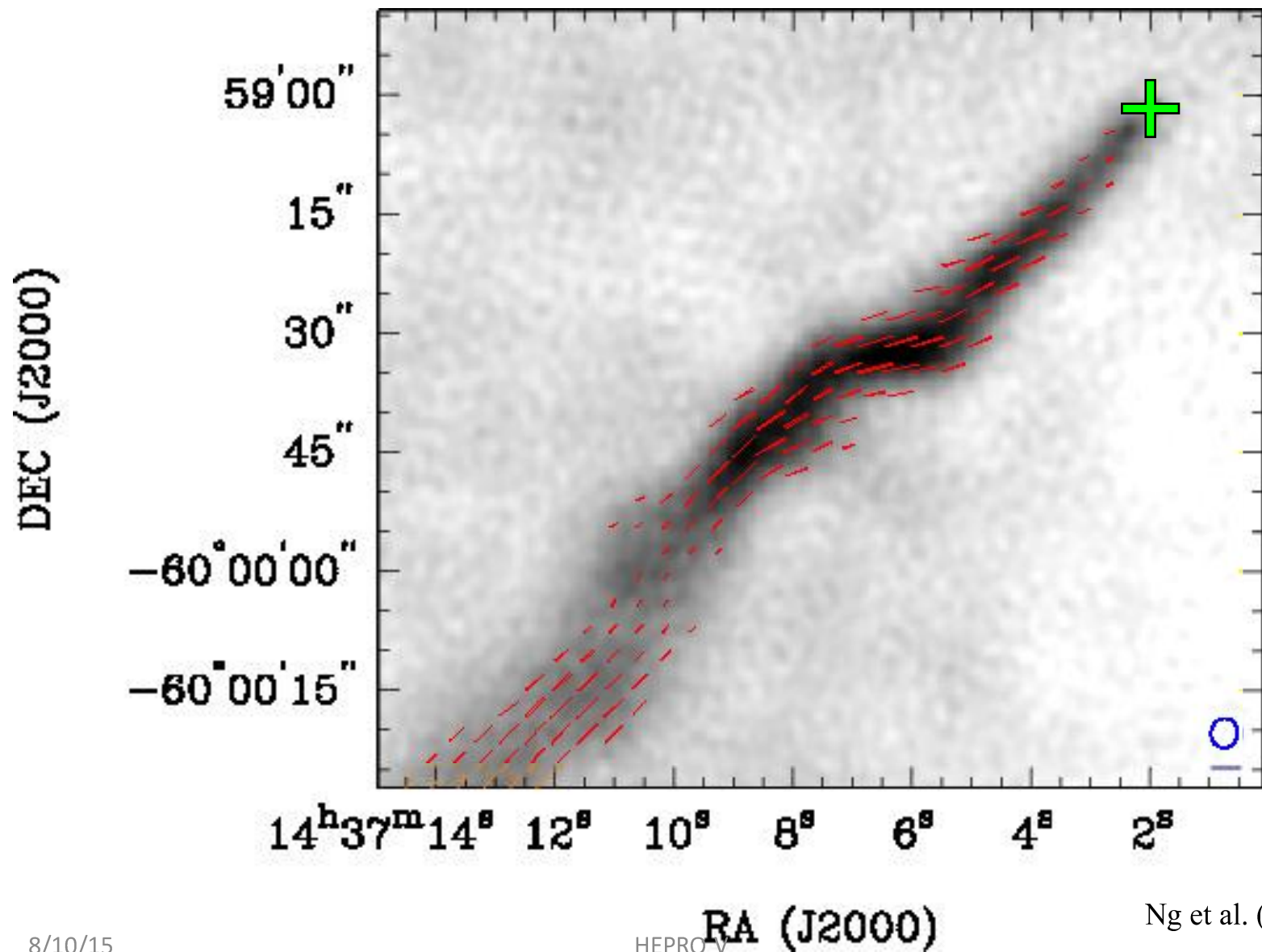
$$\Delta\chi = \frac{e^3}{8\pi^2\epsilon_0 m_e^2 c^3} \lambda^2 \int_0^d n_e(s) B_{\parallel} ds \equiv \text{RM} \lambda^2$$

G315.9–0.0: the Frying Pan



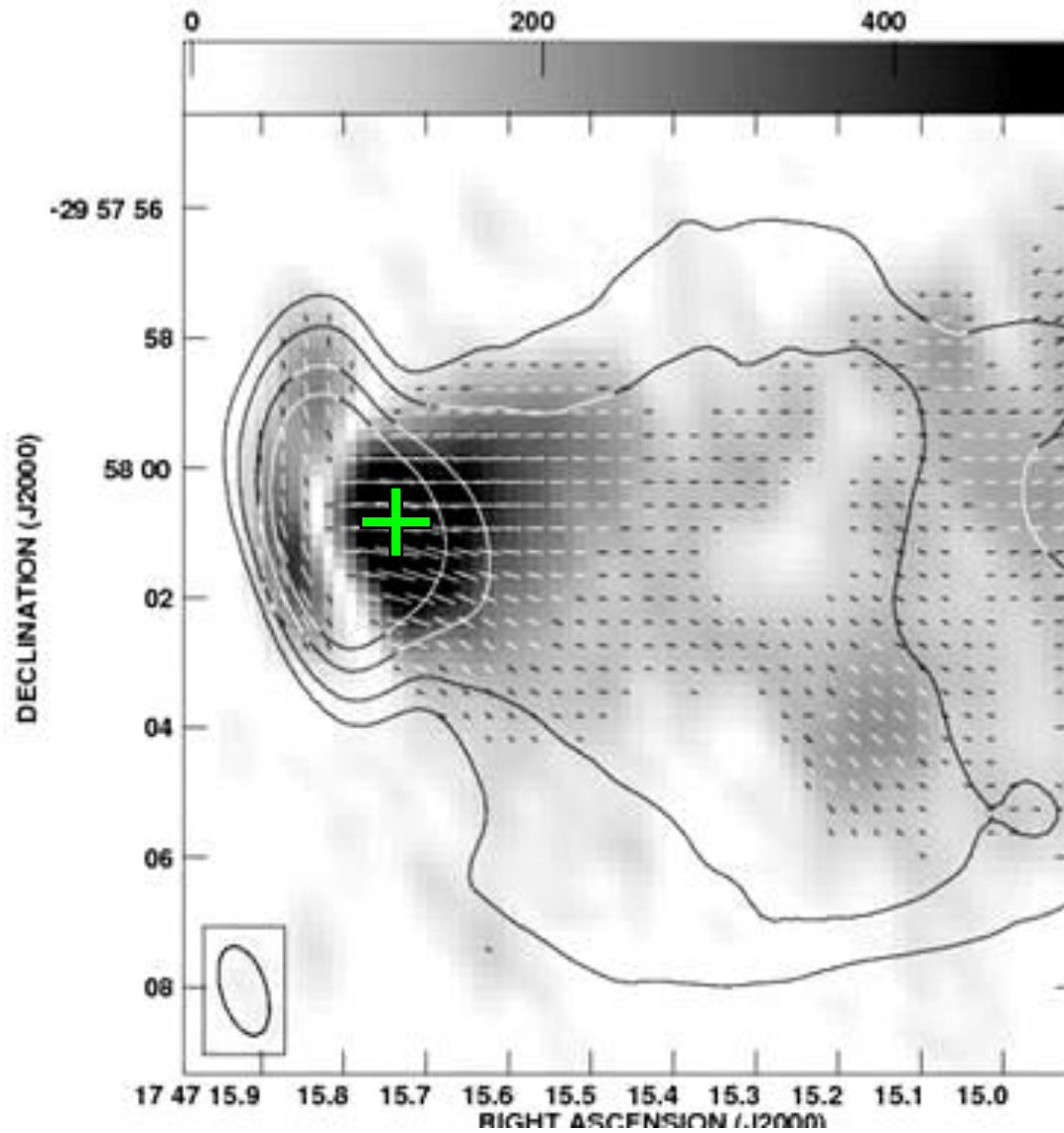
Ng et al. (2012)

ATCA 6cm

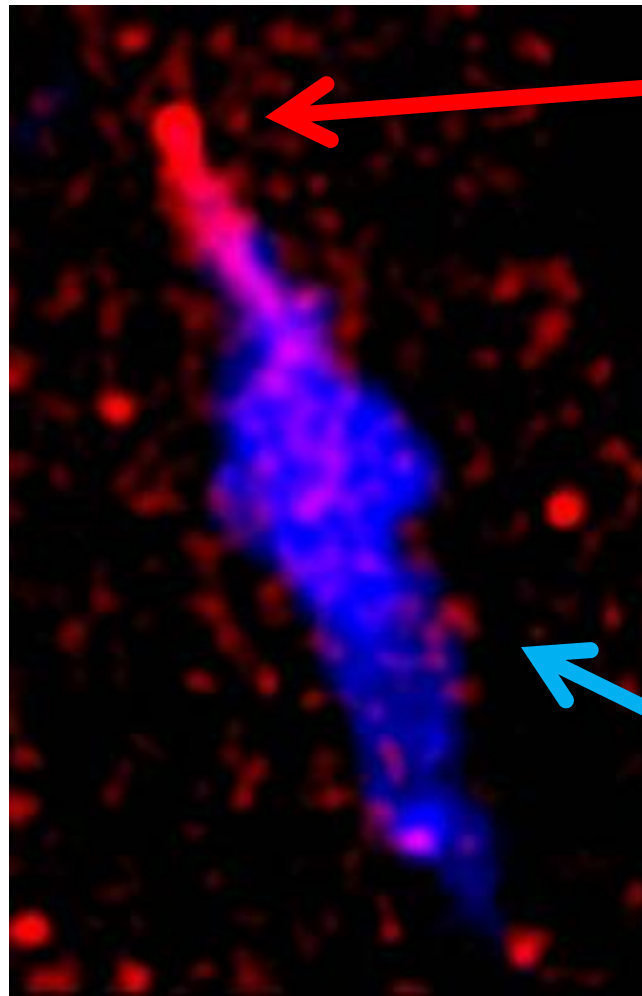


Ng et al. (2012)

The Mouse / PSR J1747-2958

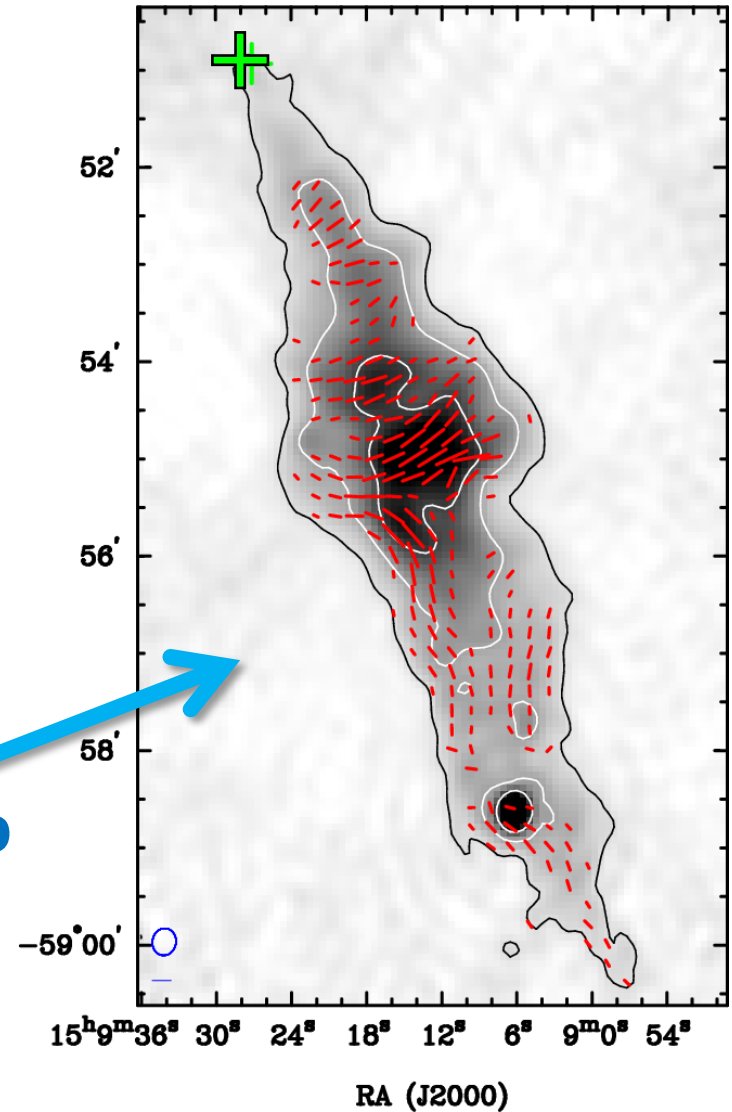


G319.9-0.7 / PSR J1509-5850



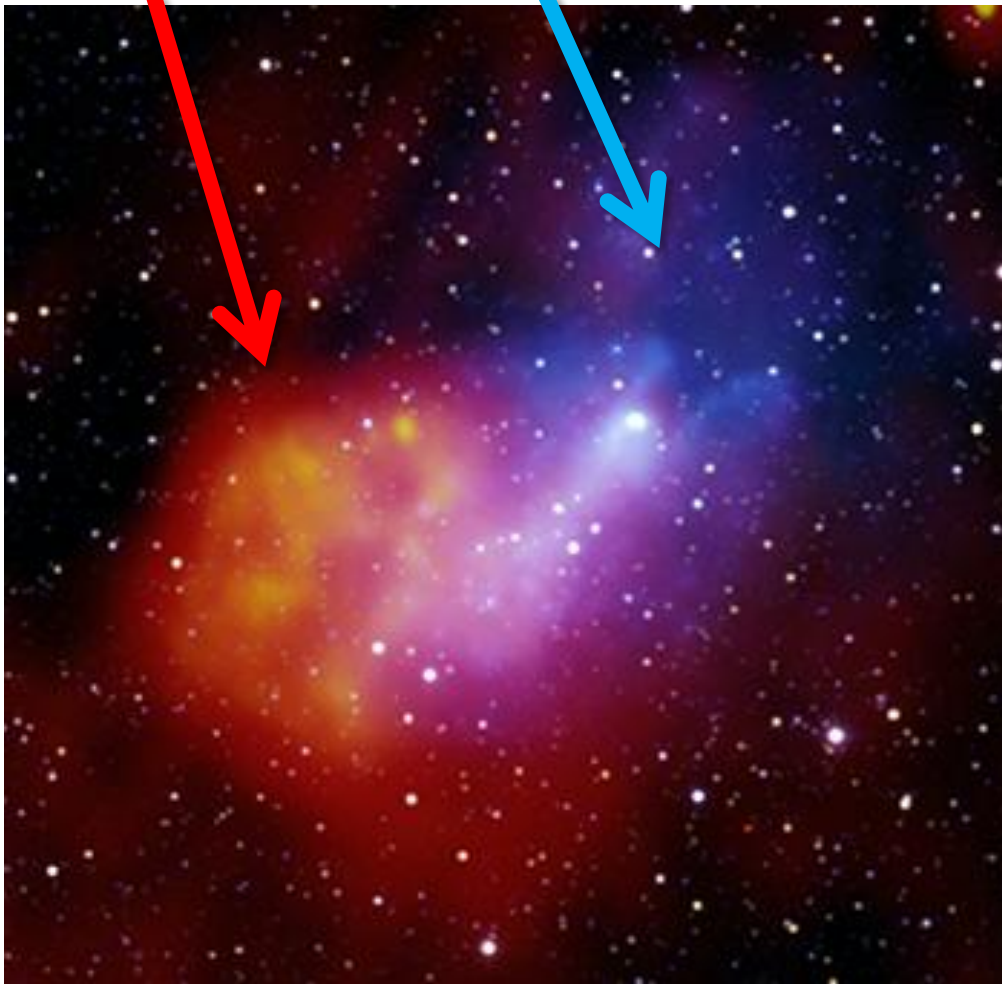
X-ray

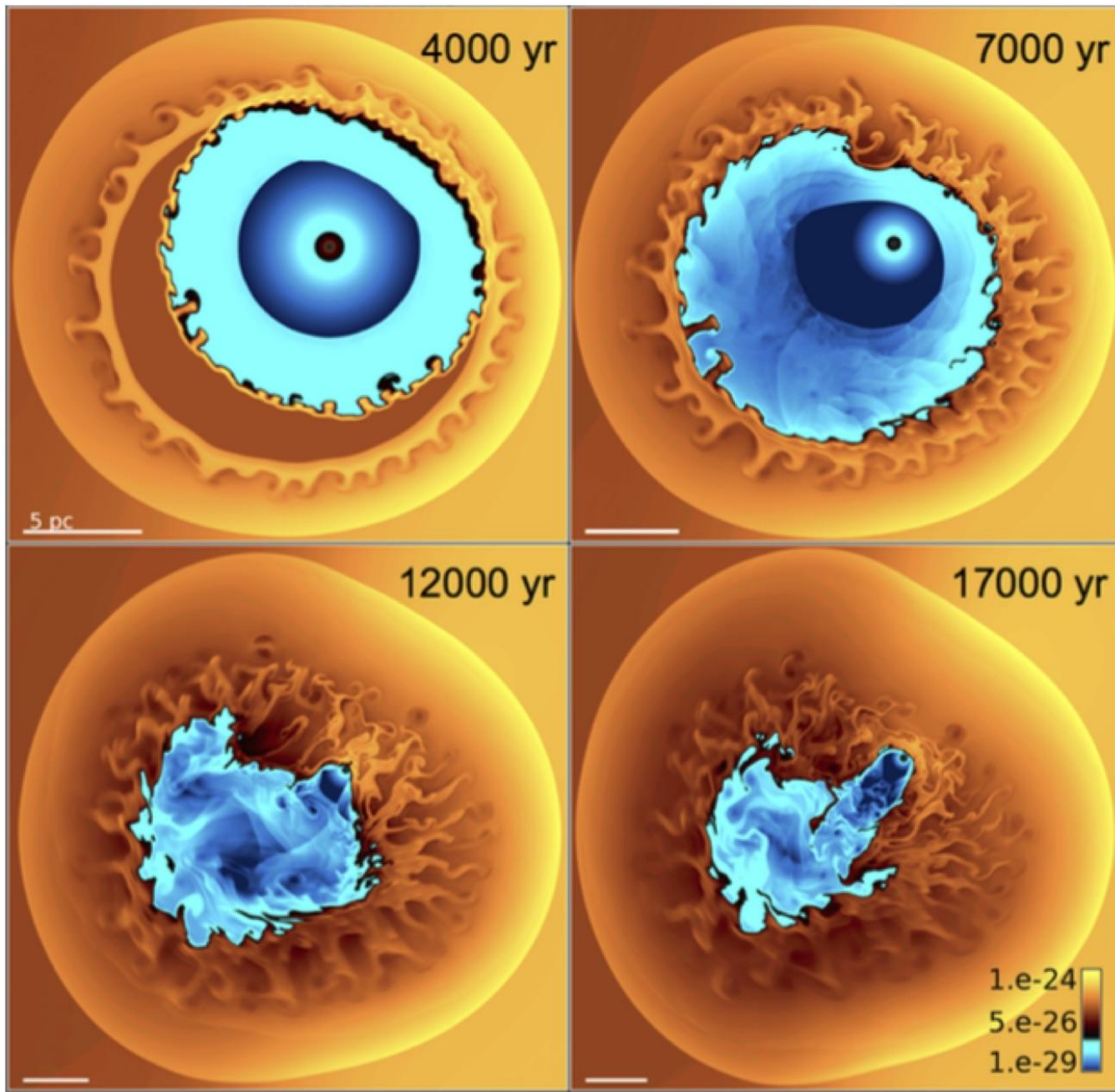
Radio



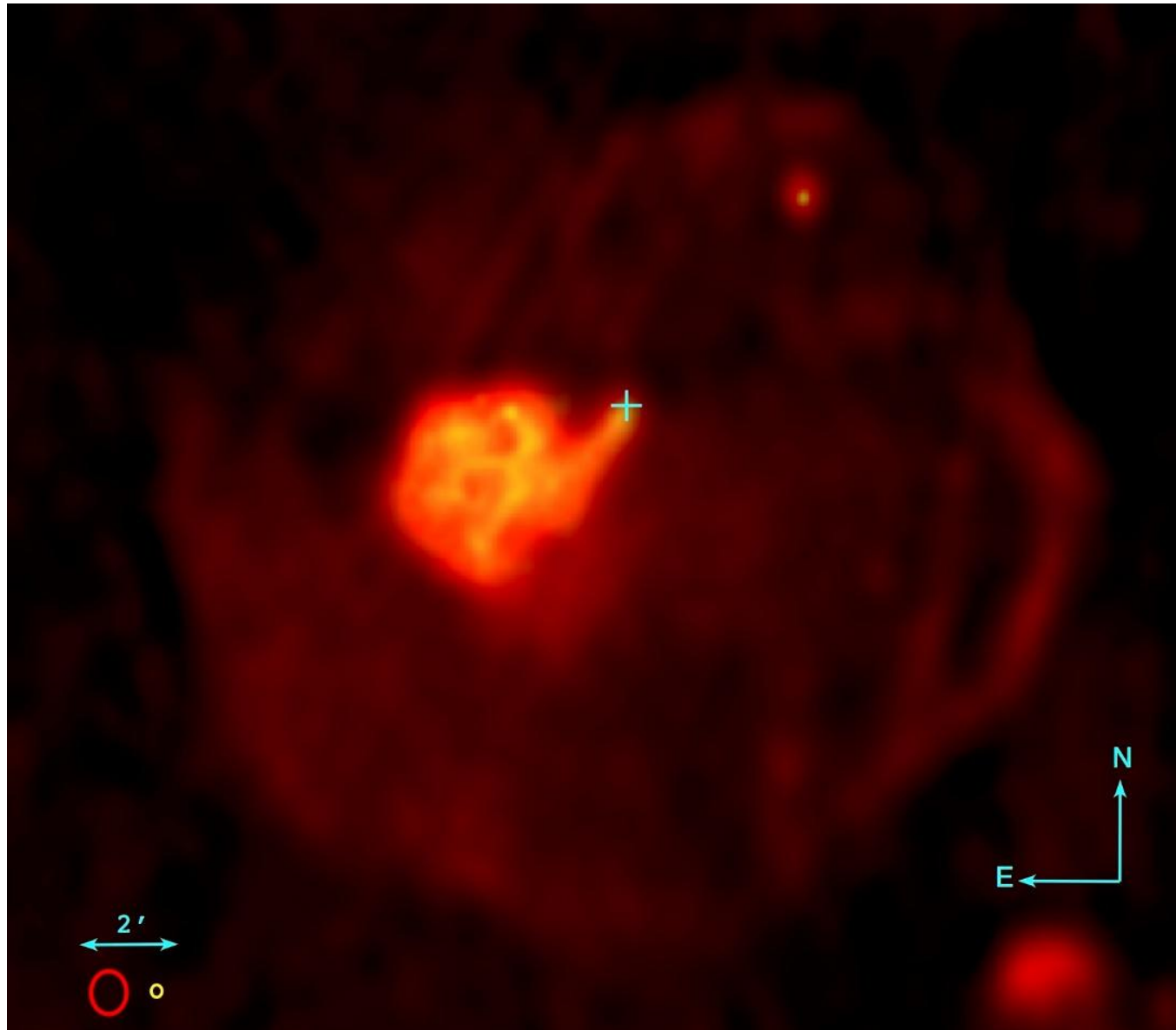
The Snail G327.1-1.1

ATCA 6cm *Chandra X-ray*





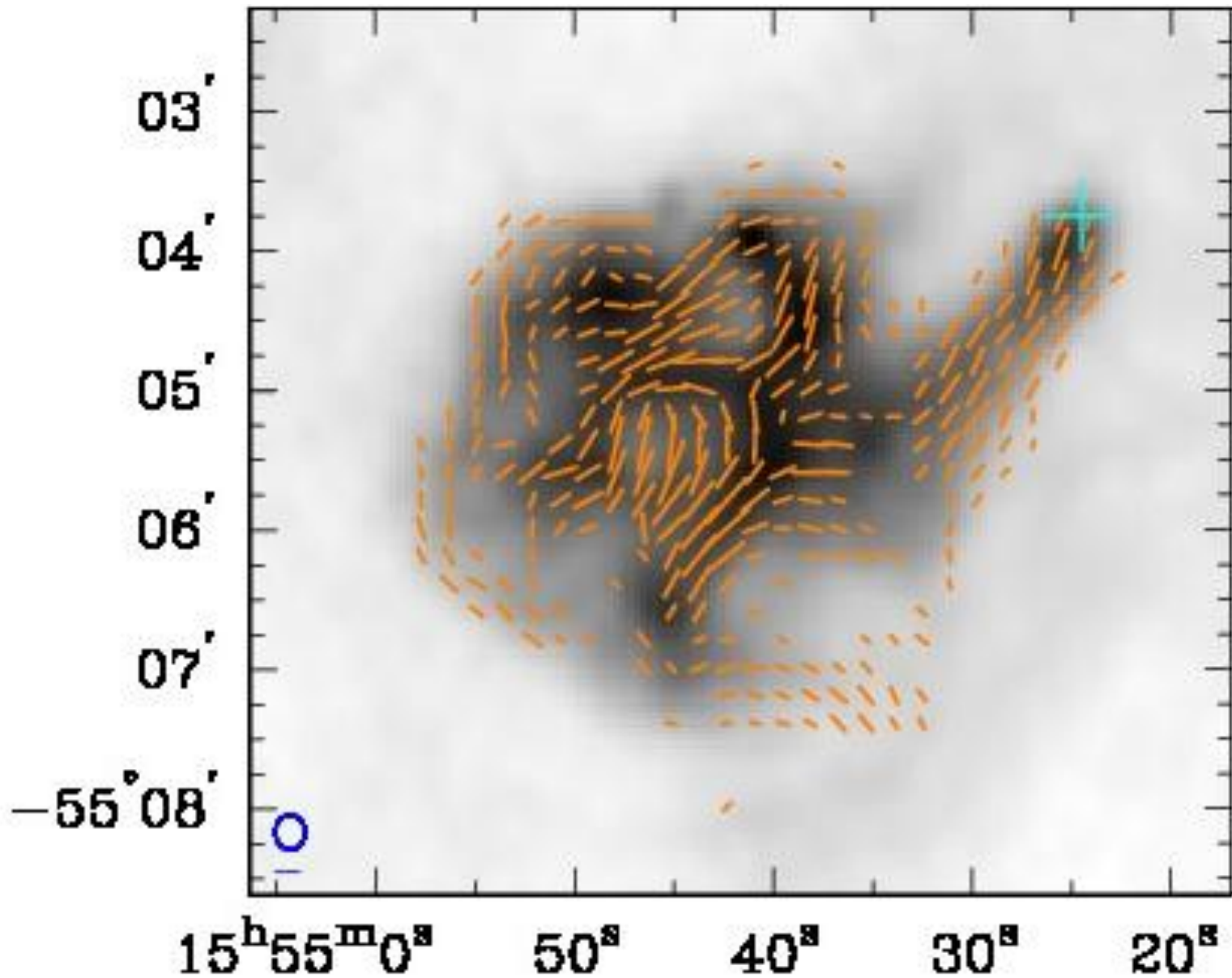
The Snail G327.1-1.1



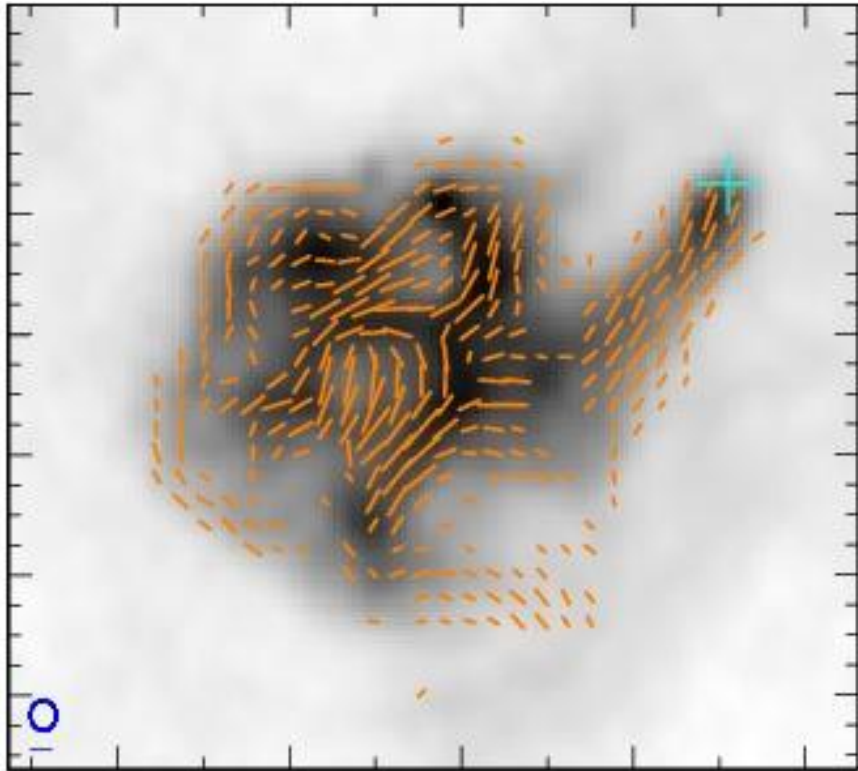
8/10/15

Ma, Ng, et al. (2015)

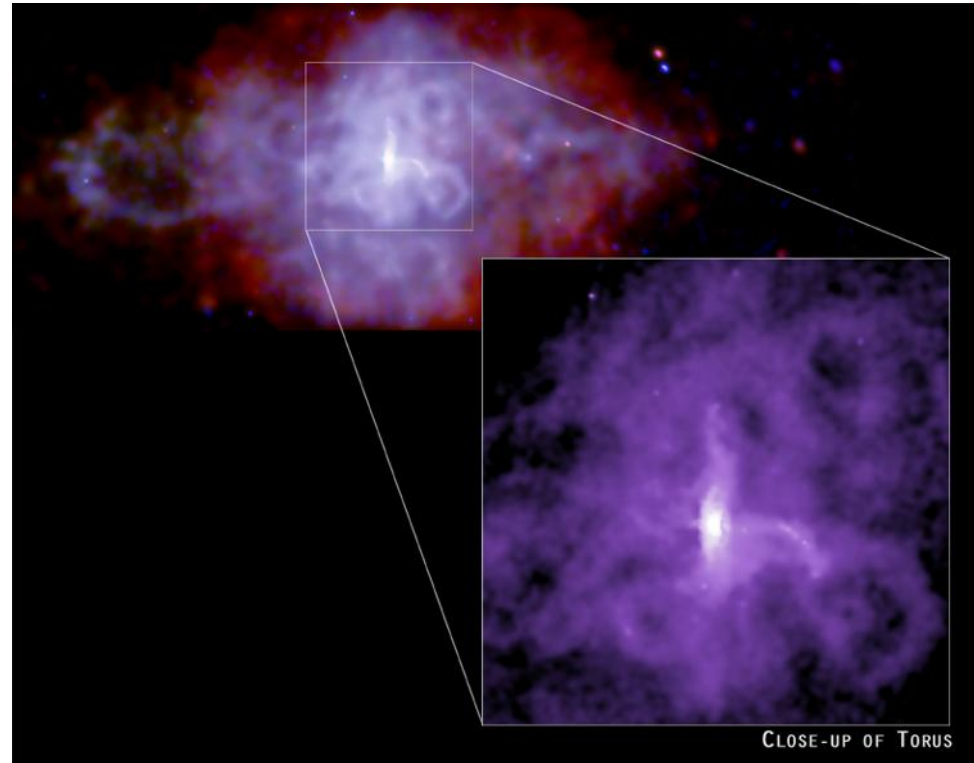
The Snail G327.1-1.1



Magnetic Loops



Ma, Ng, et al. (2015)



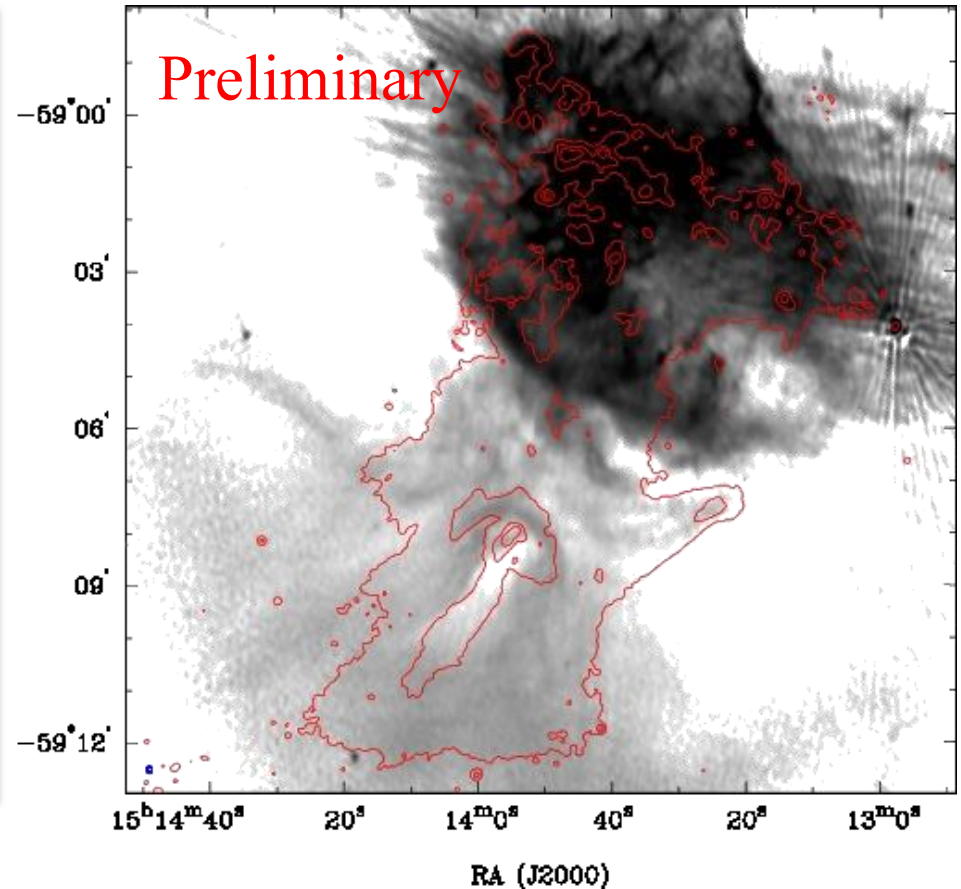
Slane et al. (2004)

The Cosmic Hand: MSH 15-52

X-ray (*Chandra*)



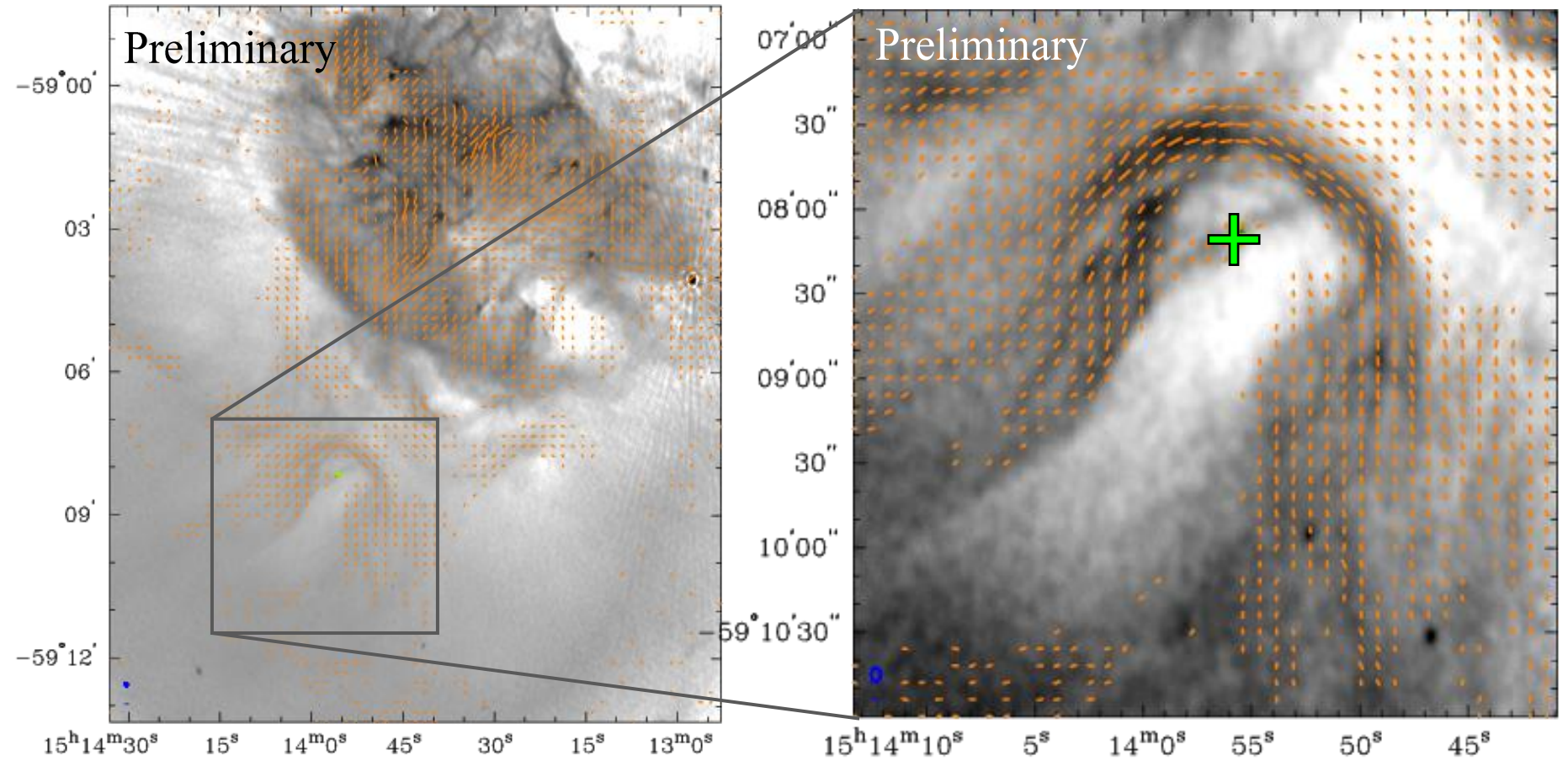
Radio 6cm (ATCA)



NASA/CXC/SAO/P.Slane, et al.

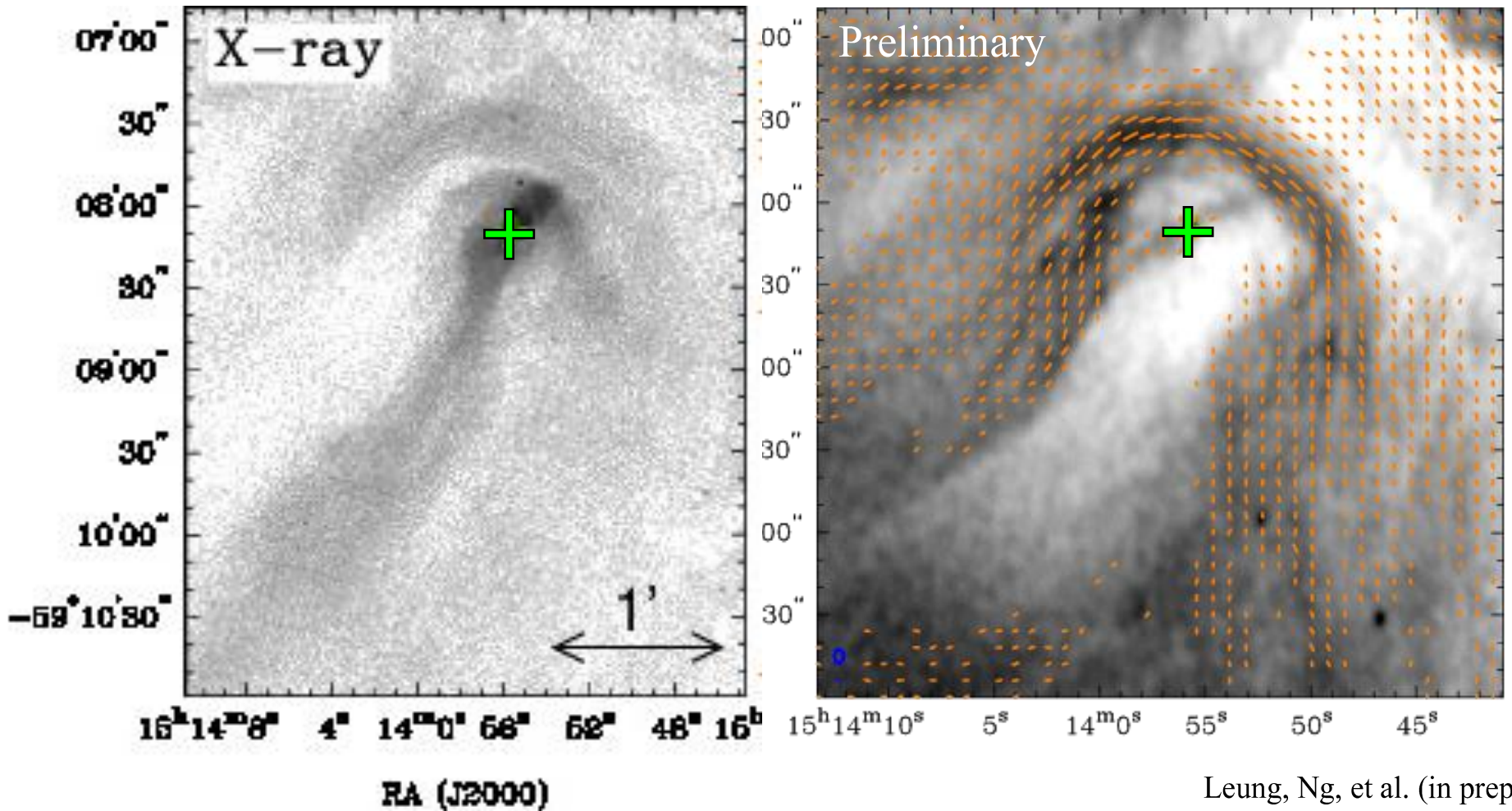
Leung, Ng, et al. (in prep.)

The Cosmic Hand: MSH 15-52



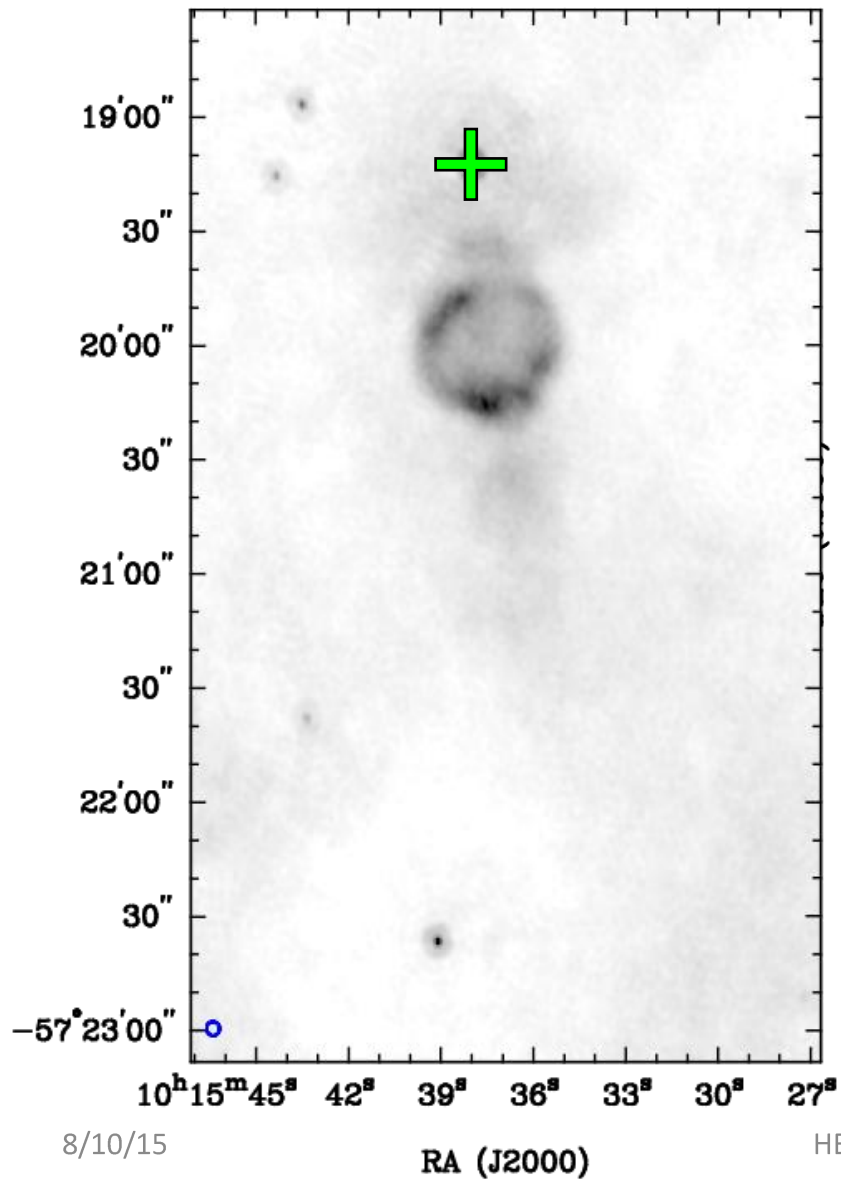
Leung, Ng, et al. (in prep.)

The Cosmic Hand: MSH 15-52

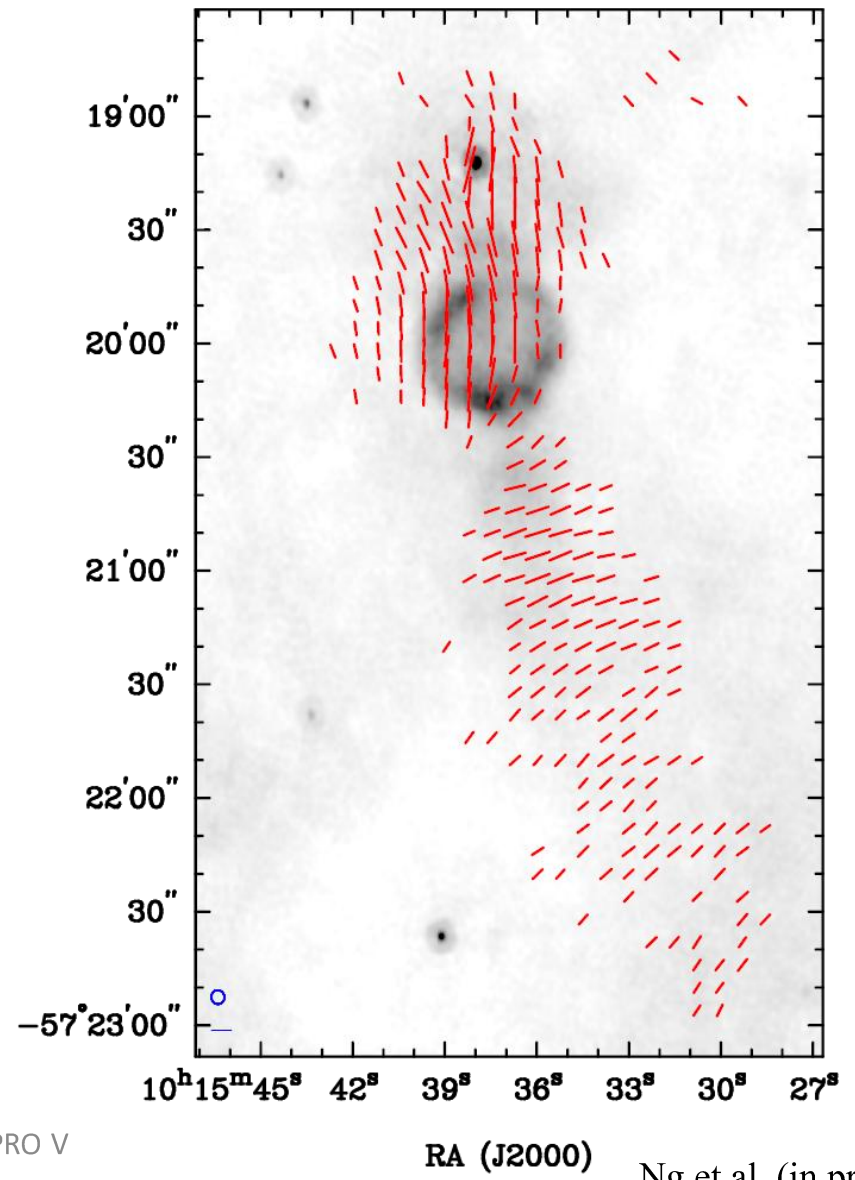


PSR J1015-5719

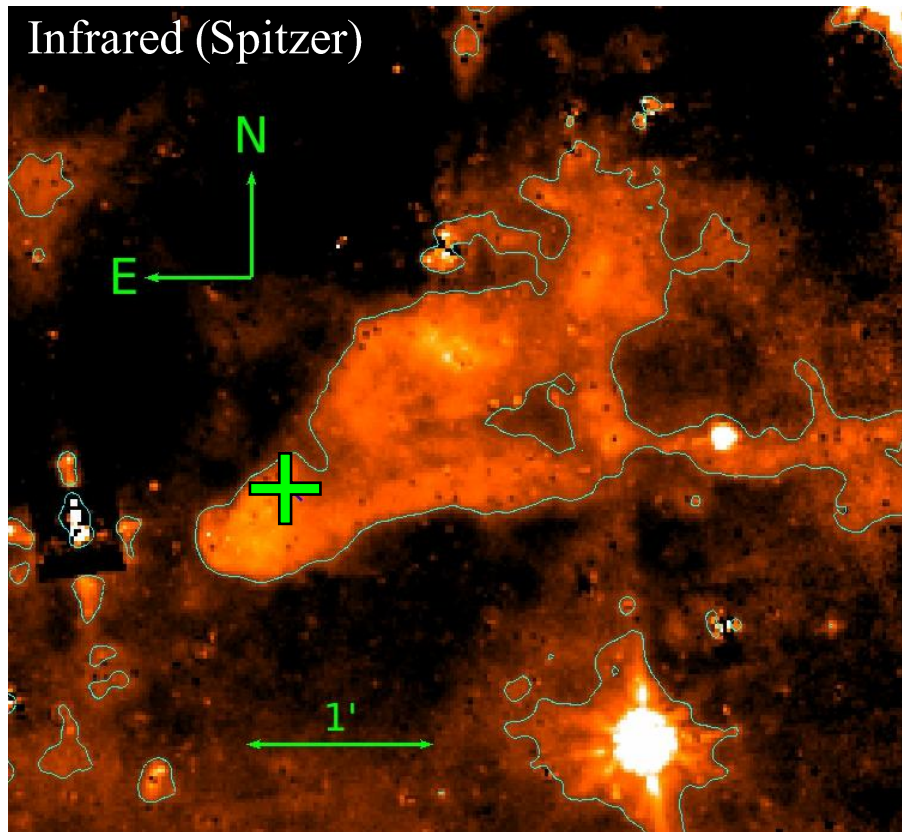
Radio 6cm (ATCA)



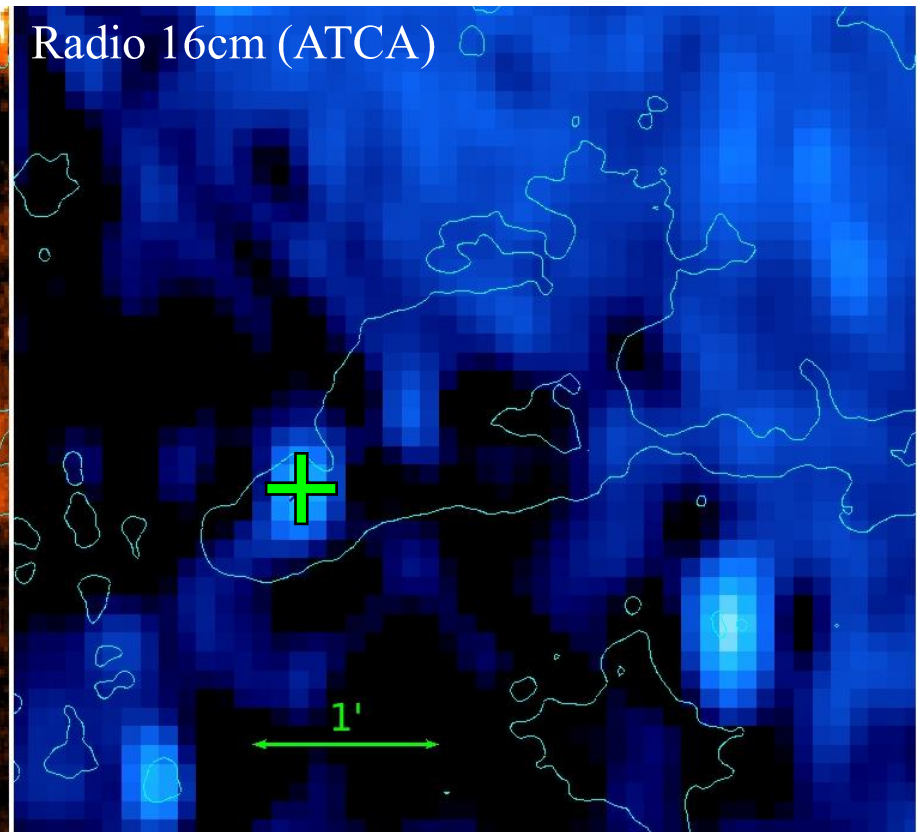
B-field configuration



IR Bow-shock PWN J1549-4848?

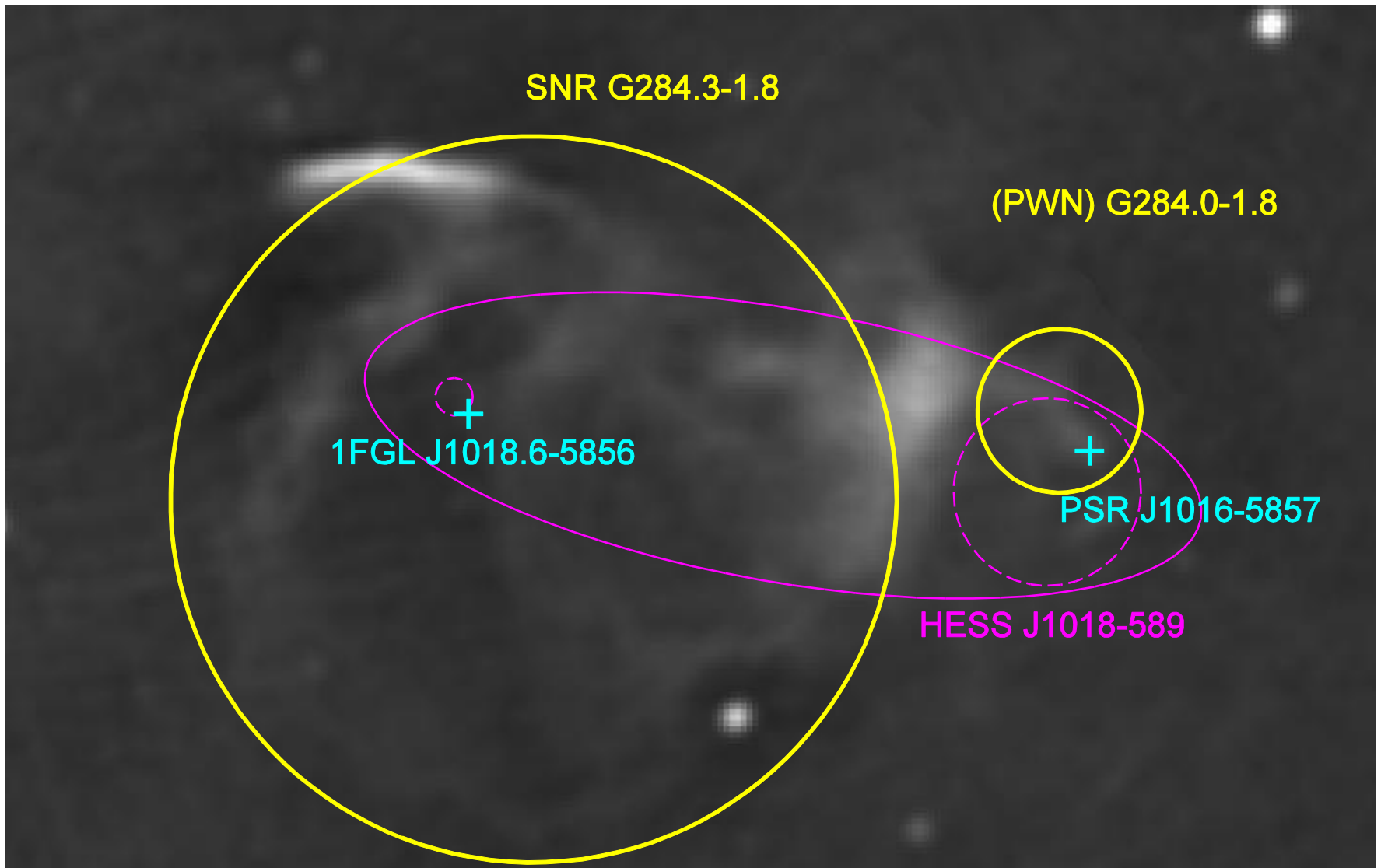


Wang et al. (2013)

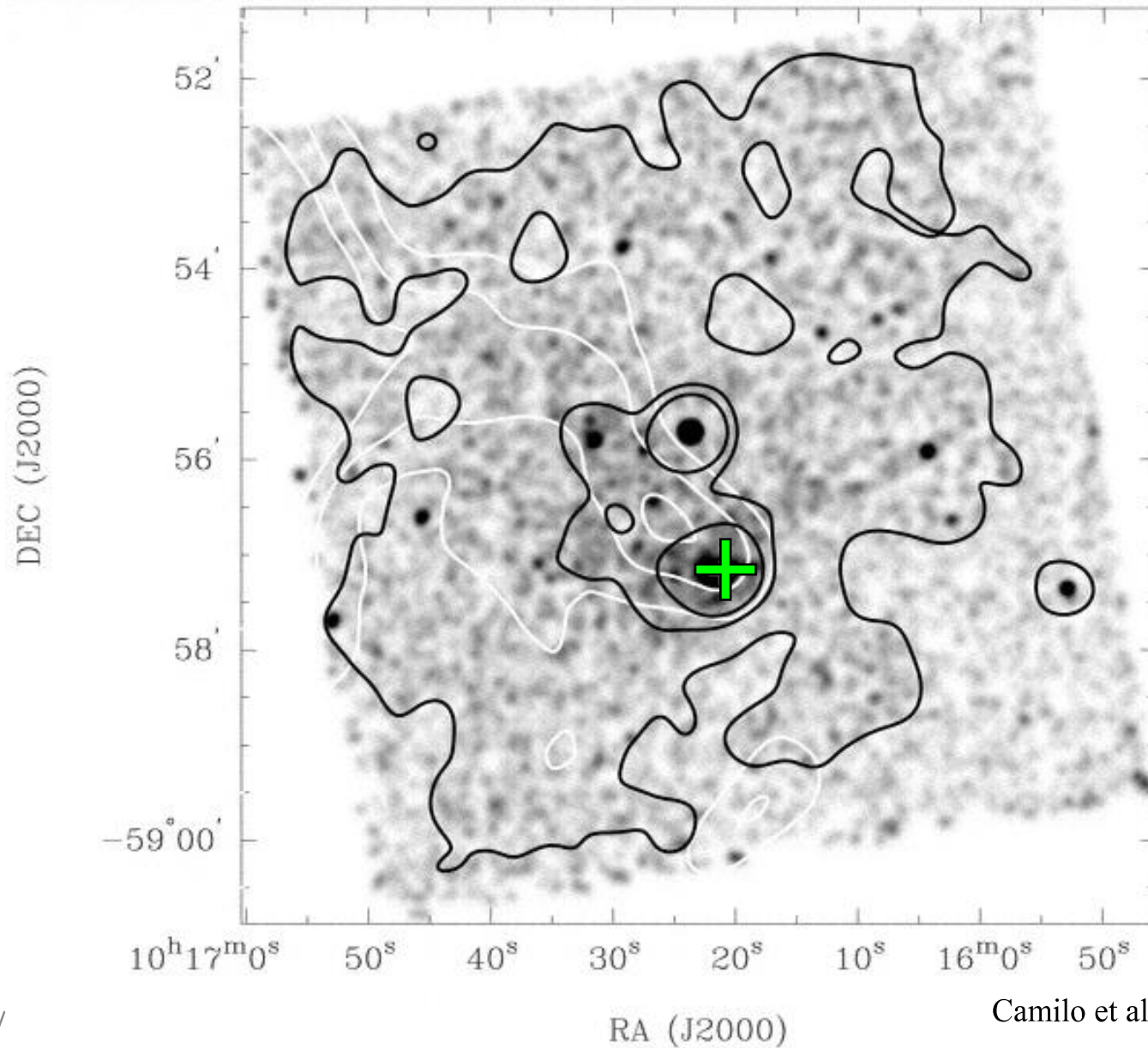


Ng et al. (in prep.)

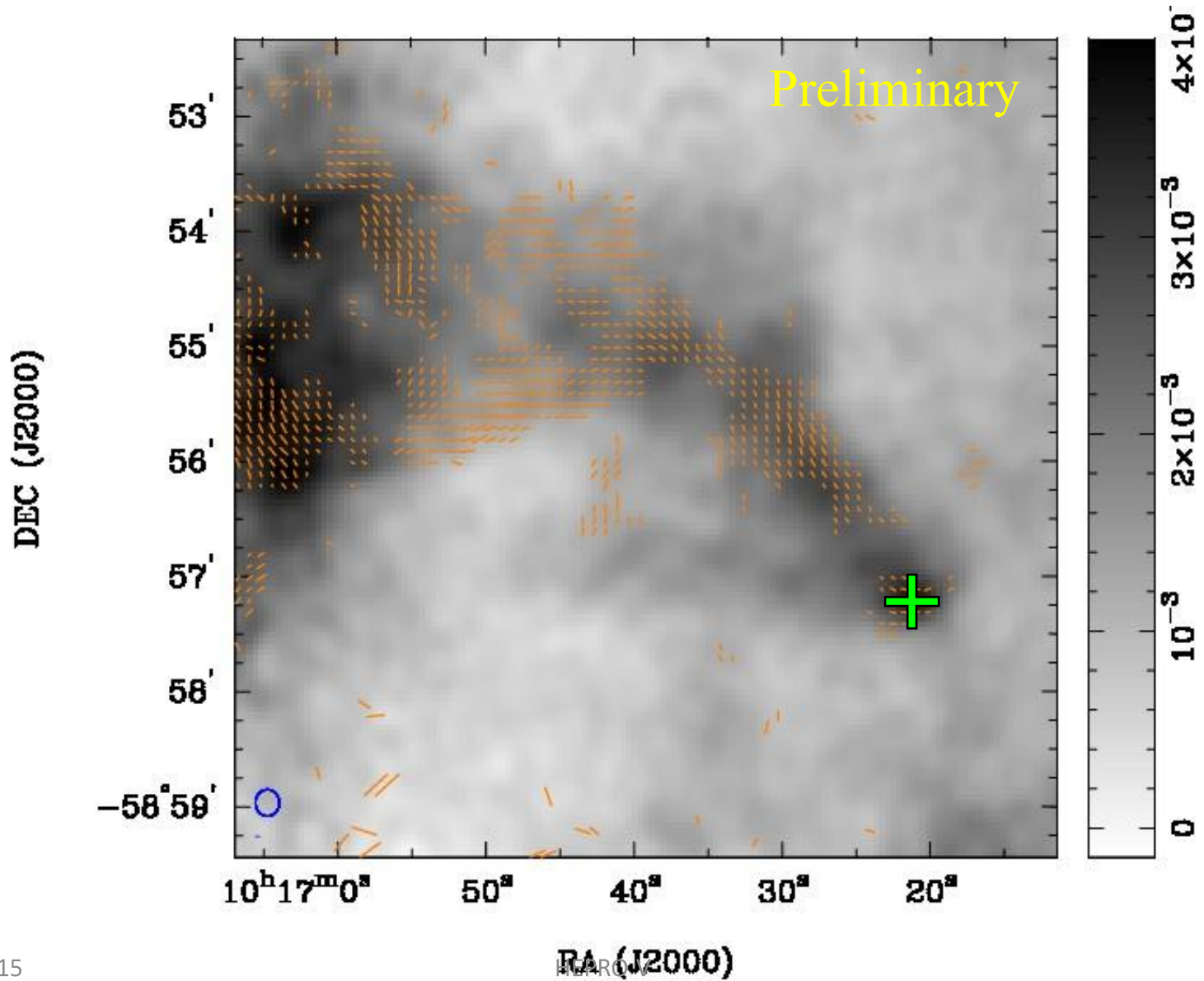
G284.0-1.8/PSR J1016-5857



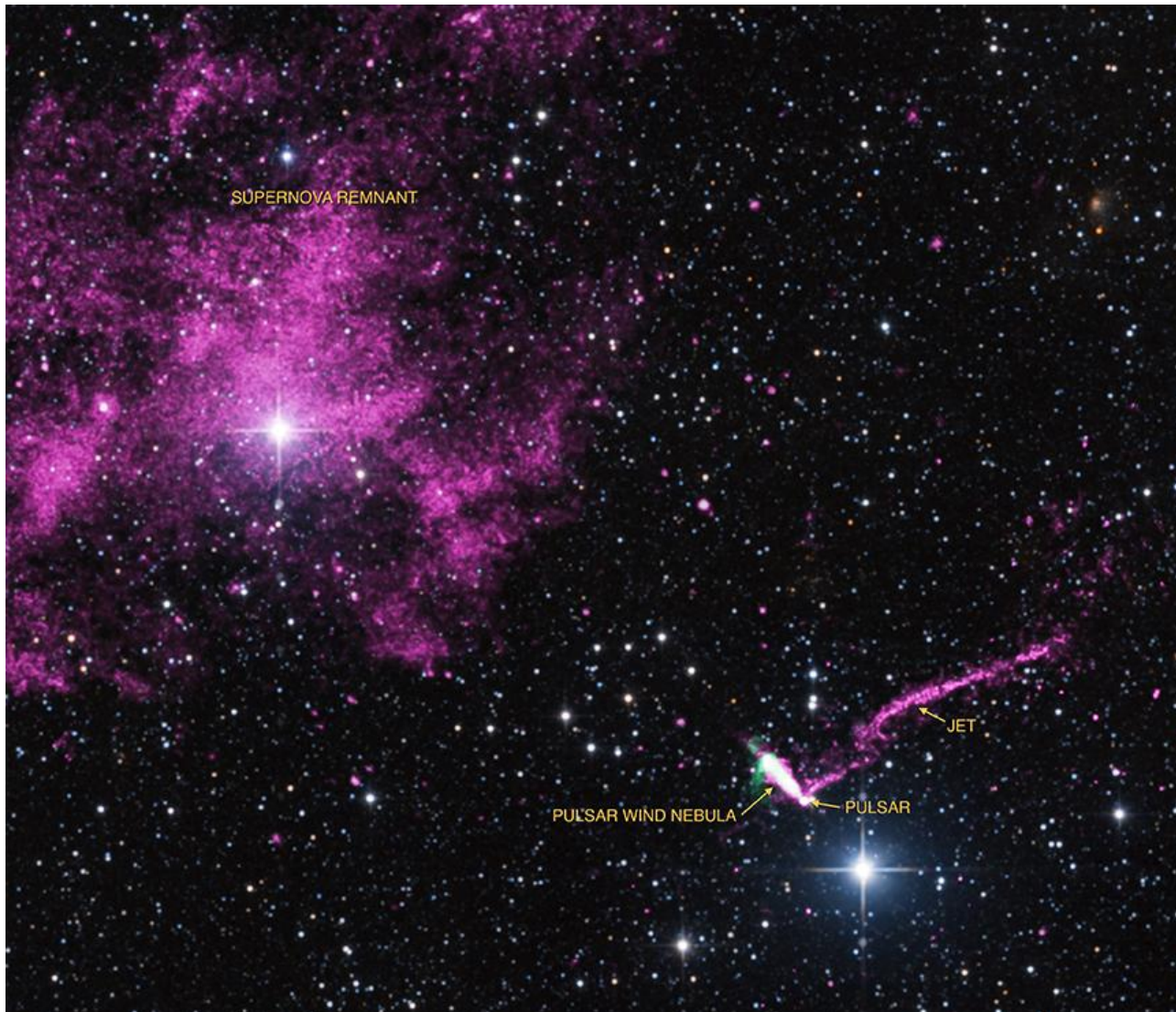
G284.0-1.8 X-ray (*Chandra*)



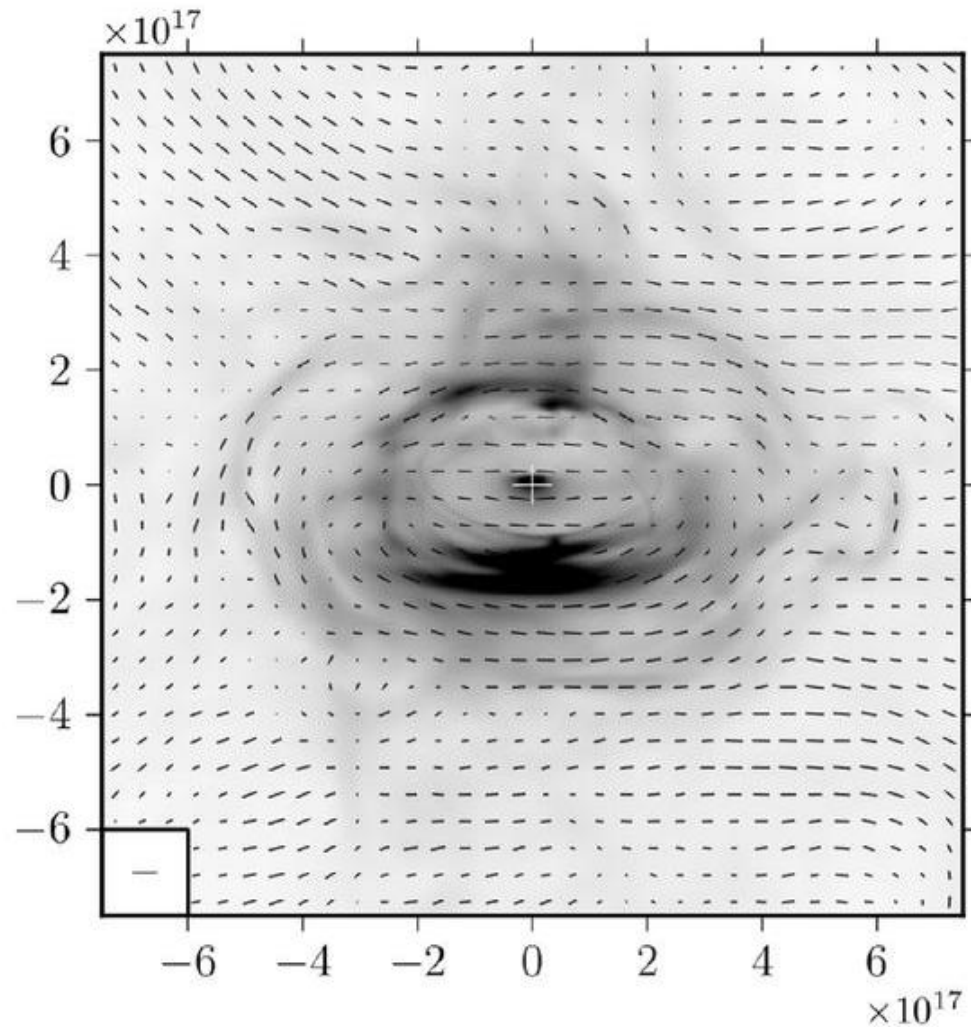
G284.0-1.8 ATCA 6cm



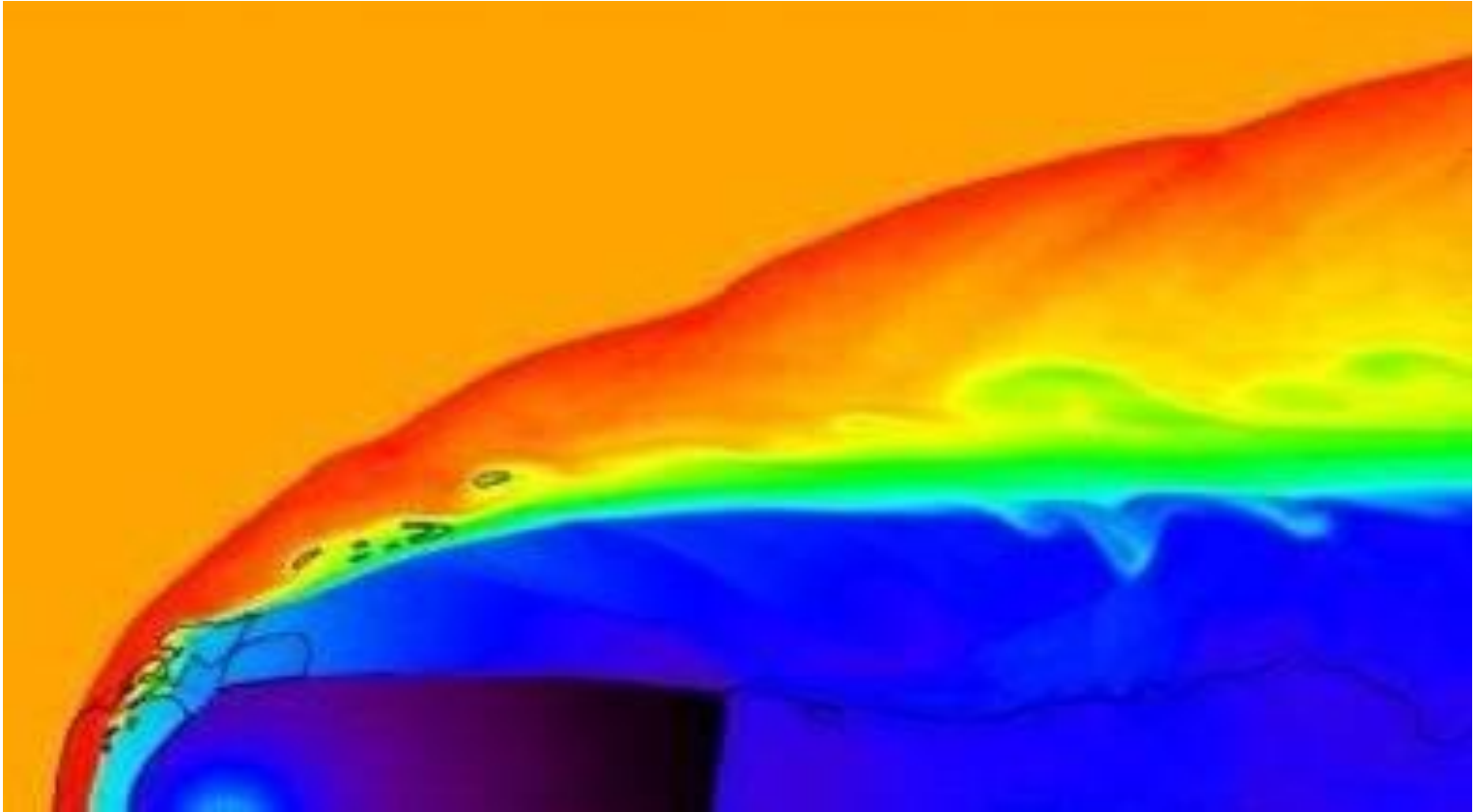
Lighthouse Nebula/IGR J11014-6103



Modeling

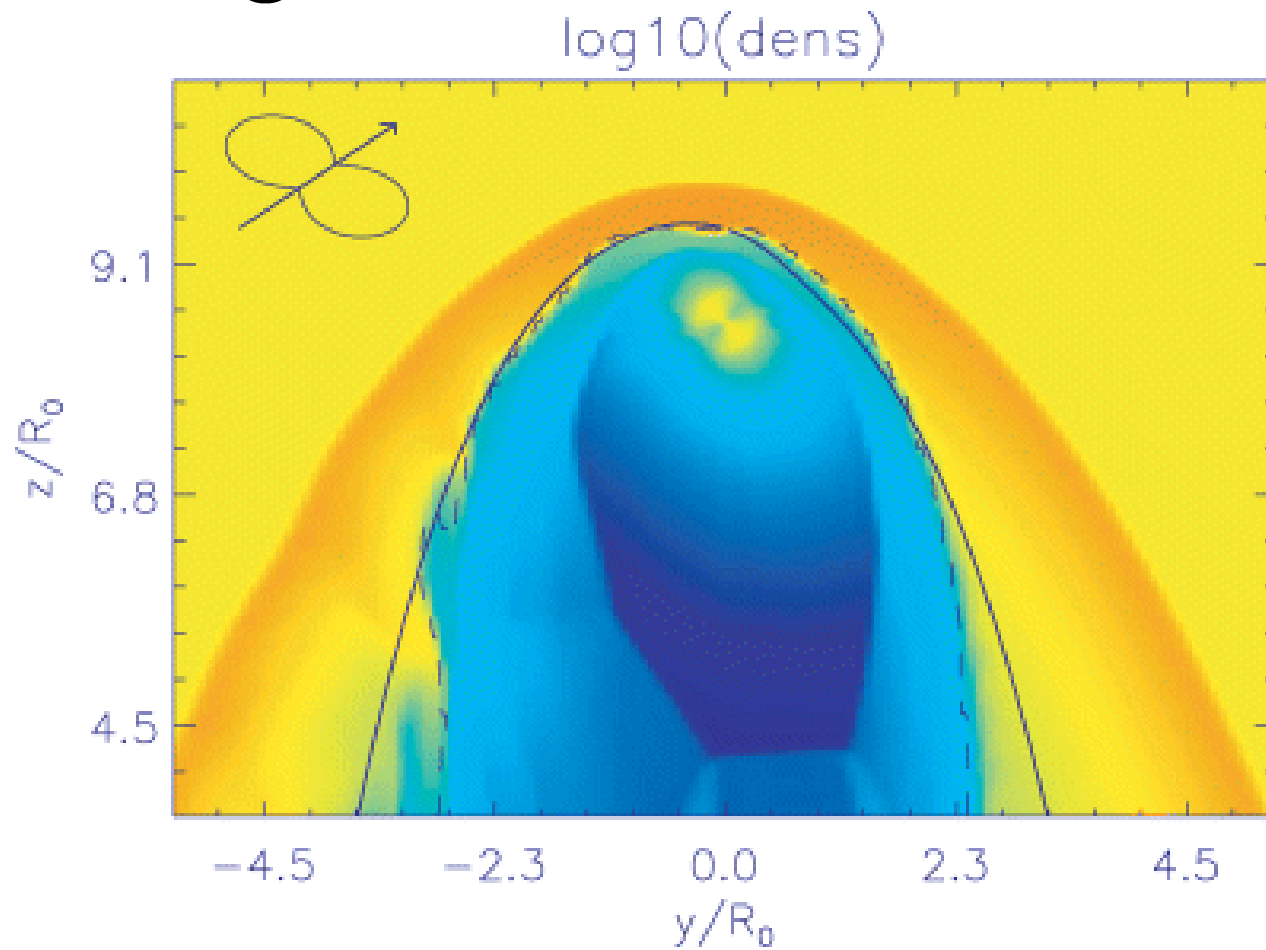


Modeling



Bucciantini et al. (2005)

Modeling



Vigelius et al. (2007)

Summary

- Radio polarization measurement as a powerful tool
- Large diversity in field geometry:
 - Parallel / perpendicular B -field in bow-shocks
 - Highly ordered field in crushed PWNe
 - Filamentary structure in young objects
- Theoretical modeling and MHD simulations needed:
 - Connection with other physical parameters, e.g. Mach number, pulsar spin orientation.
 - Particle acceleration and transport